

# VLBI with NOEMA (**N**orthern **E**xtended **M**illimeter **A**rray)

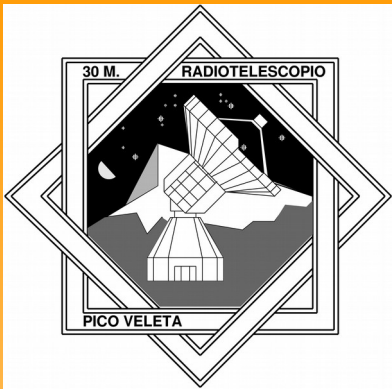
*Michael Bremer*

*Institut de Radio Astronomie Millimétrique (IRAM)*

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*Royal Observatory Madrid, February 8<sup>th</sup> - 9<sup>th</sup> 2016*

- The NOEMA extension – current status and future plans



# Geographical Overview

IRAM supports two mm/sub-mm instruments in central Europe, connecting baseline length: 1146 km



# NOEMA: 15-m aperture dishes on railway tracks

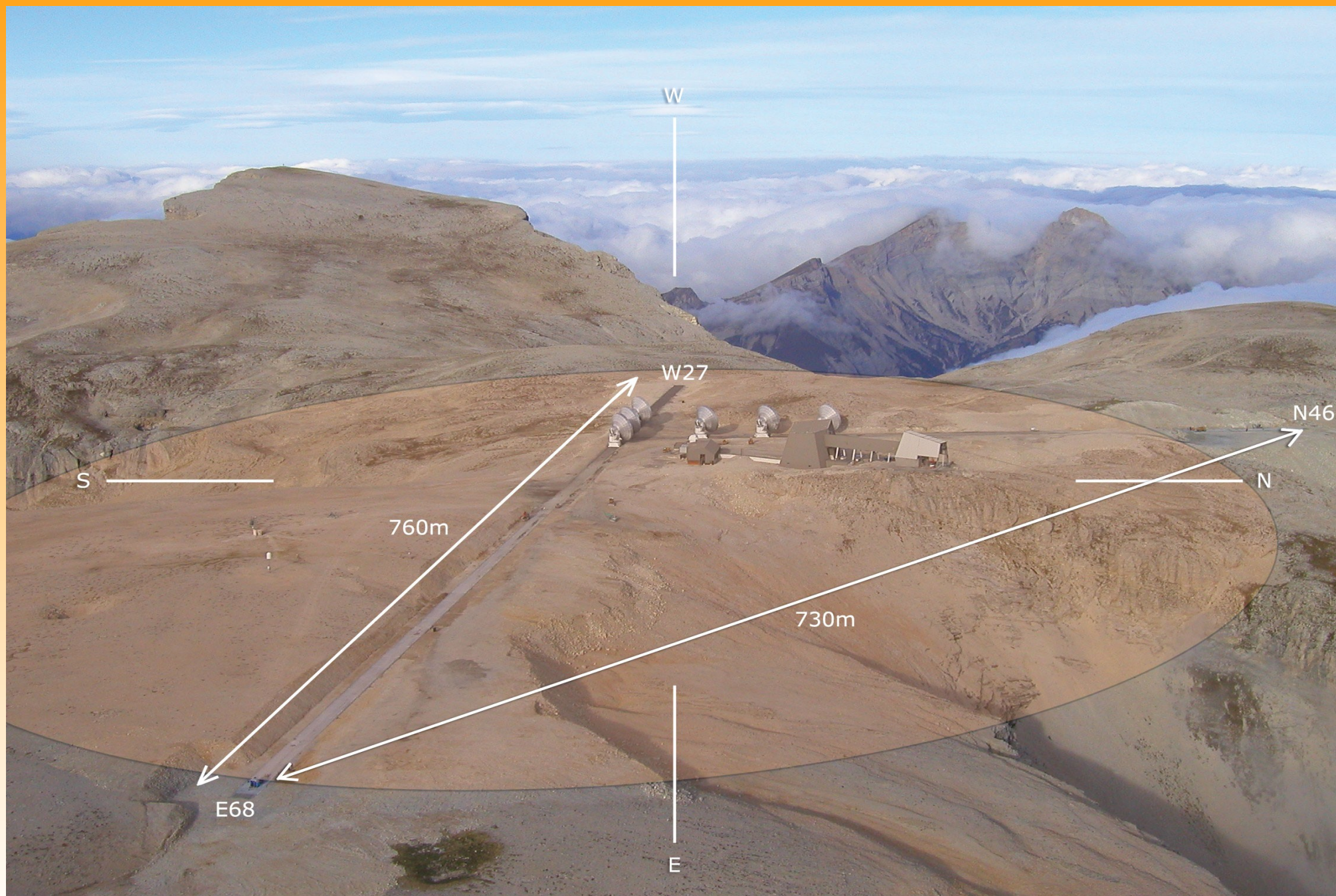


Plateau de Bure

Hautes Alpes, France

Altitude 2550m

# NOEMA: synthesis array with two tracks (E-W,N-S)



# NOEMA 2016: SSB receiver tuning ranges

	Band 1	Band 2	Band 3	Band 4
RF range*/[GHz]	80–116	129–174	201–267	277–371
T <sub>rec</sub> /[K] LSB	40–55	30–50	40–60	30–50
T <sub>rec</sub> /[K] USB	40–55	40–80	50–70	30–50
G <sub>im</sub> /[dB]	-10	-12 ... -10	-12 ... -8	-20
RF LSB/[GHz]	80–104	129–165	201–264	277–359
RF USB/[GHz]	104–116	164–174	264–267	289–371

\* center of the 4.2-7.8 GHz IF band;

**All bands are in dual polarization, Single-Sideband, 3.6 GHz/polarization**

**Filter wheels:** Currently we have lambda/4 plates installed for band 1 and band 3

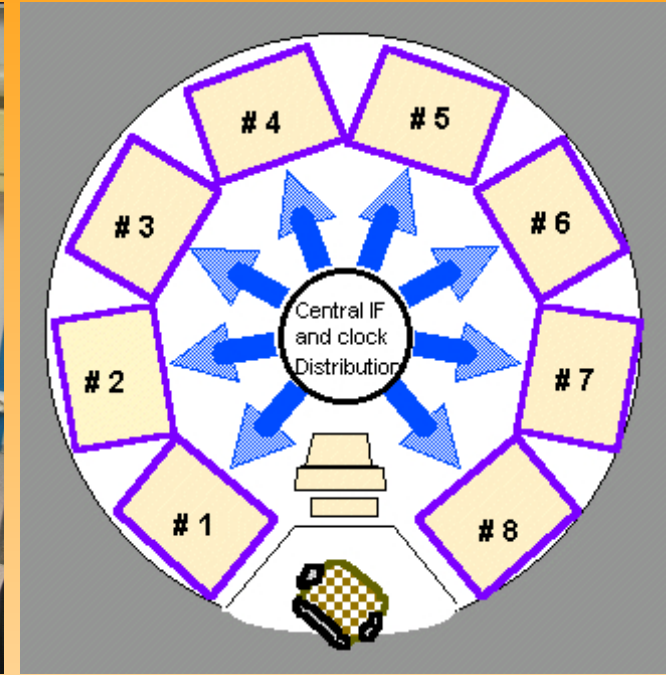
# Present NOEMA VLBI equipment



Swiss EFOS-38 maser (built in 2006),  
with low phase noise quartz



Mark4 Formatter +  
Mark5A recorder



Narrow-band correlator (In  
operation since 2000)



Rohde & Schwarz SMA100 B22  
frequency generator

Only the Narrow-band correlator is VLBI capable,  
not the more recent continuum WIDEX backend.

The phased array operation is currently  
limited to 1 Gbit/sec data rates ,  
VLBI correlation at MPIfR in zoom mode.

# That looks already nice. Why do we want to upgrade the Interferometer further?

The main reason is ALMA (Chajnantor, Chile).

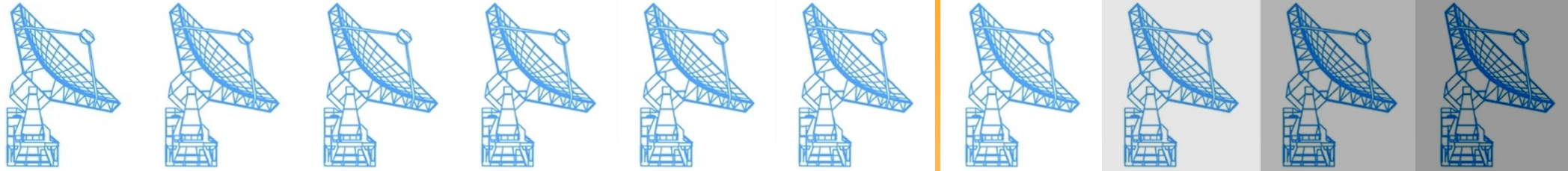
The idea is to have a Northern instrument that is not too far behind ALMA, and that is a bit easier to get access to.

Also, part of the Northern sky is not accessible for ALMA.

Take the old six-antenna Plateau de Bure array and add:

- More collecting area, more simultaneous baselines
- More receiver bandwidth
- Longer baselines

# NOEMA extension progress: February 2016



6-antenna PdBI (since 16-Dec-2001)

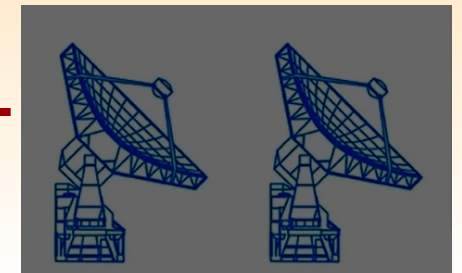
NOEMA antenna 7 (inaugurated 2014)

NOEMA antenna 8 (near completion)

NOEMA antenna 9 (construction started)

NOEMA antenna 10 (funded)

NOEMA antennas 11+12 (with more partners?)





# NOEMA extension progress: February 2016

## **Receivers:**

Progressive conversion from SSB (2x4 GHz) to 2SB (4x8 GHz) receivers, 3 of 7 antennas are today equipped but those 3 are bandpass-filtered to stay compatible with the others (*observing continues during upgrade*).  
→ 8 antennas, each with 2SB receivers, foreseen for end of 2016

## **Backend:**

Polyfix correlator installation is planned to start in late 2016

## **Baselines:**

Track construction will start in 2016: additional service tracks first, extension of the main E-W track in 2017

# NOEMA

## Northern Extended Millimeter Array:

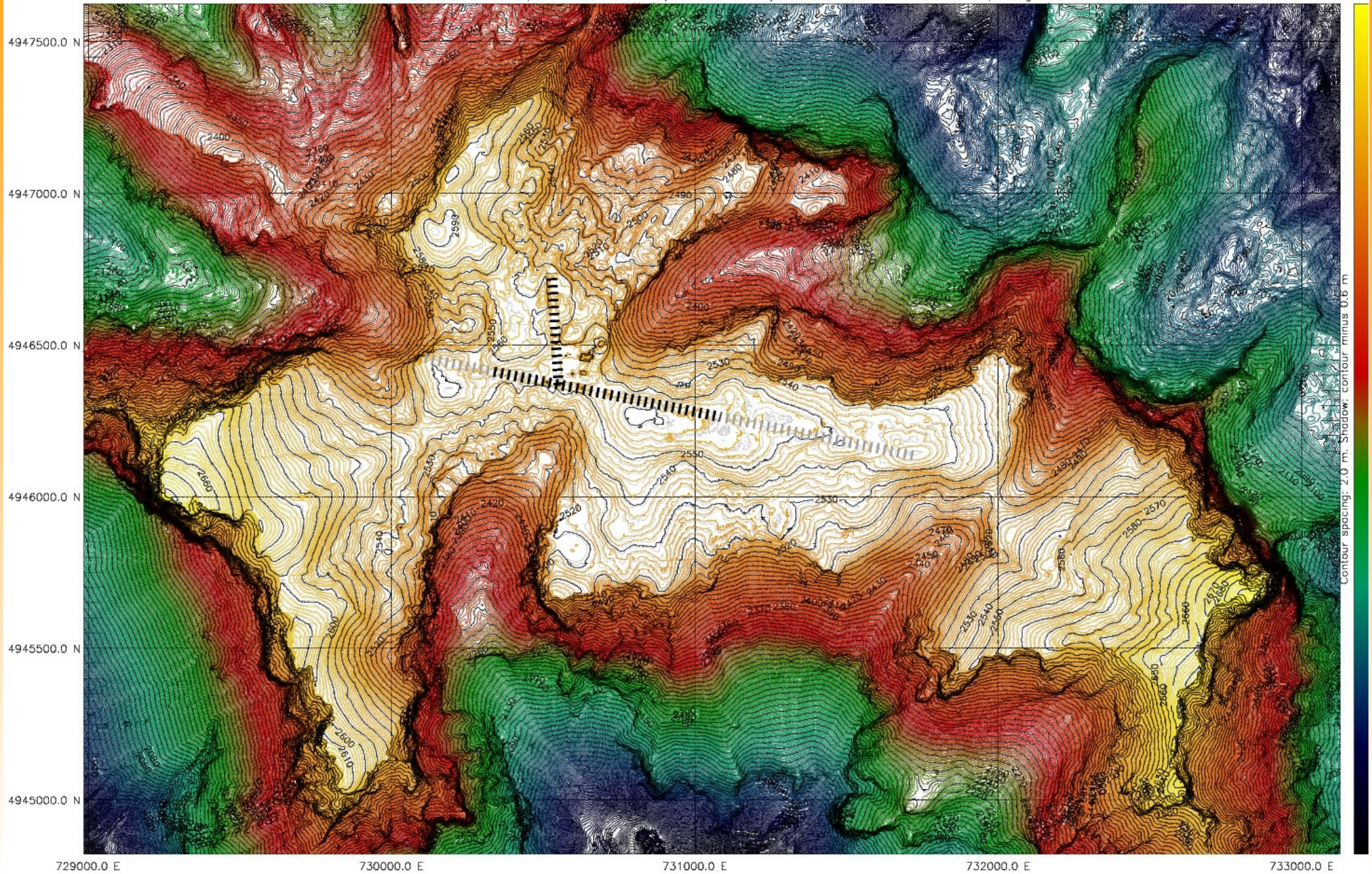
	PdBI	NOEMA now	NOEMA
antennas	6	7	10 (12 ?)
baselines	15	21	45 (66 ?)
bandwidth	2 x 3.6 GHz	2 x 3.6 GHz	2 x 16 GHz
max. baseline	760 m	760 m	1600 m

**Why 12 antennas and 1600 m baselines and not more ?**

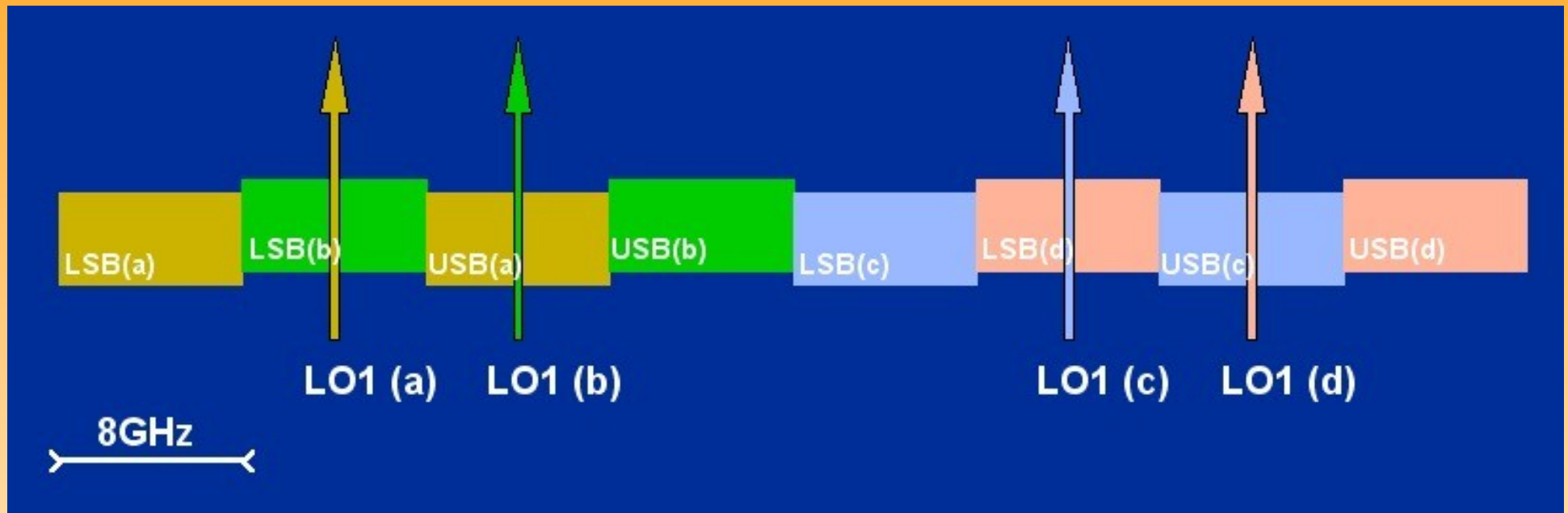


# Well, that is all the space we have ...

Plateau de Bure, aerial map. WGS 1984, UTM (Zone 31 North) All units in meters, contour spacing 2.0 m.



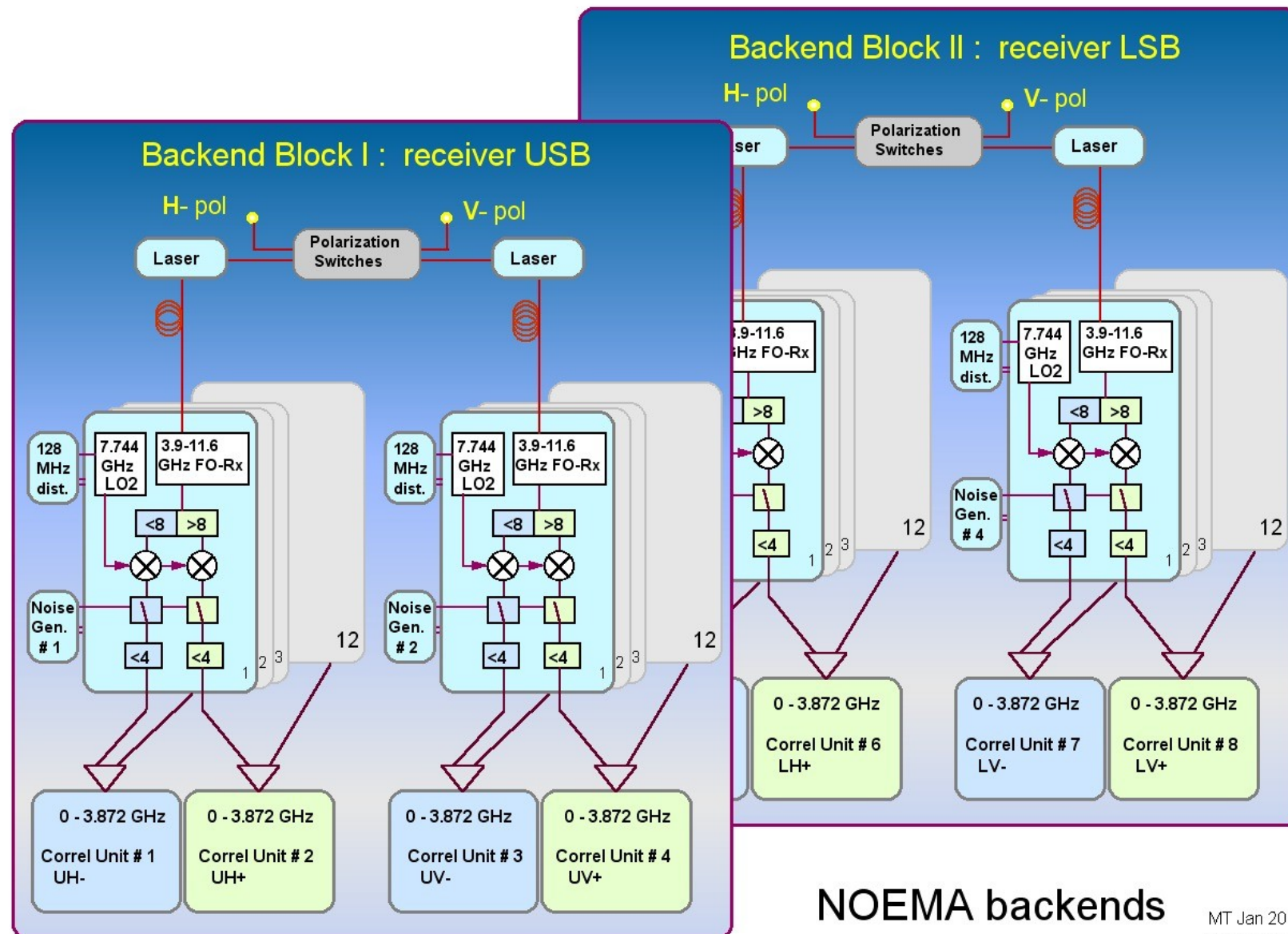
# End 2016: The 5<sup>th</sup> generation correlator on Bure (first phase – extension to VLBI capabilities later)



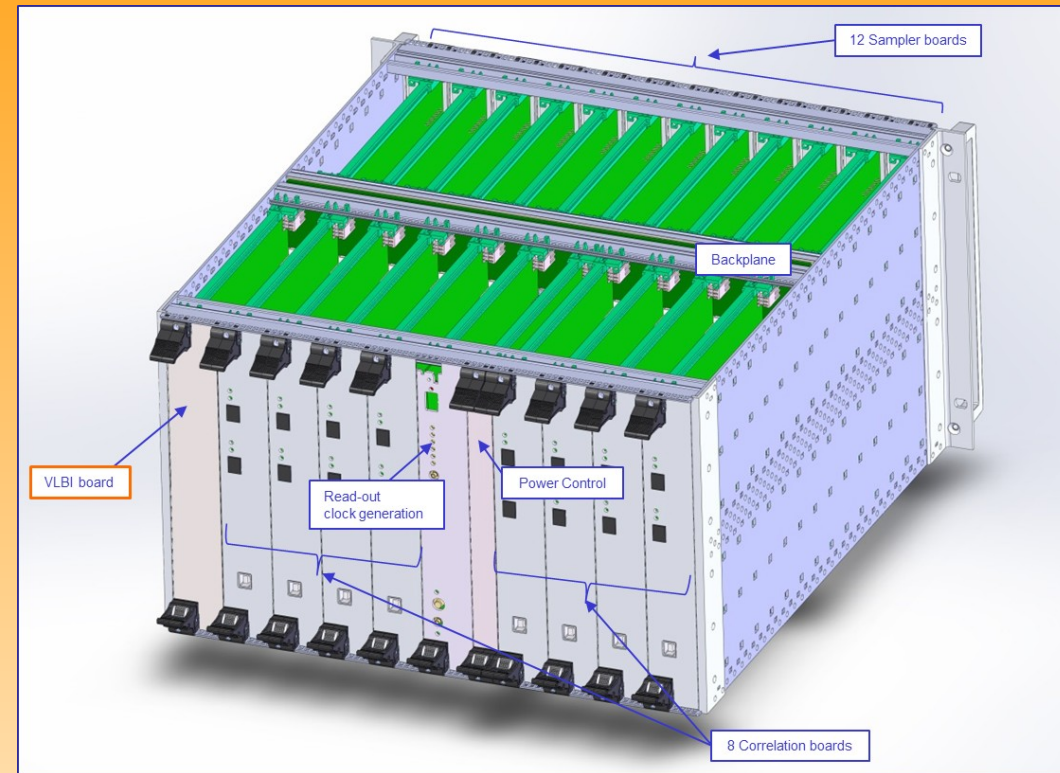
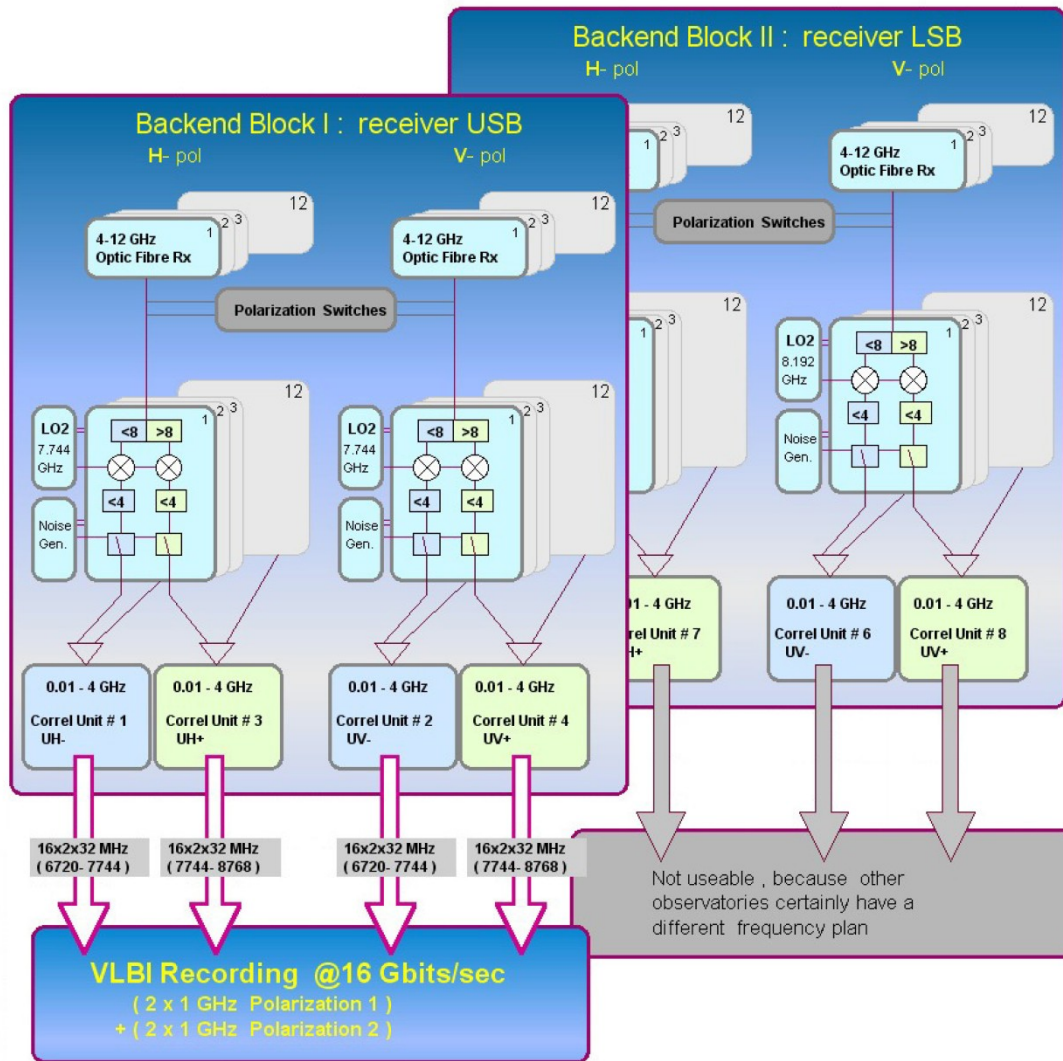
The updated NOEMA receivers will deliver 8-GHz wide sidebands in 2SB mode, dual polarization. The new correlator is based on FPGAs (Altera), highly flexible.

Architecture is VLBI compatible but will be **without** the adder cards

# End 2016: The 5<sup>th</sup> generation correlator on Bure (first phase – extension to VLBI capabilities around 2018)



# The 5<sup>th</sup> generation correlator – second phase



## VLBI mode:

16 Gbits/s for one receiver sideband  
 32 Gbits/s if the second sideband can be used

[ 64 Gbits/s are close to the design limits, and cannot be guaranteed at this time at full S/N]

# But how do we VLBI until then?

**Problem:** the old narrowband correlator will disappear at the end of 2016.

However, one single NOEMA antenna with broadband (16 Gbit/s or more) backend would be more powerful than this old phased array backend.

**Solution :** Install a broad-band VLBI backend on a single dish of the array.  
(we are working on that).

Still further in the future: a second correlator for simultaneous dual-frequency, dual-polarization observations.

IRAM has multiplied the capacities of its instruments by many orders of magnitude during its 36 years of existence. And it's still going on ...



**Thank you!**