# **EVN Amplitude Calibration**

## Jun Yang

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# **Calibration Accuracy**

## **Session 3/2012**

========	========	========	========
Station	18 cm	6 cm	5 cm
Jb2	0.08(6)	0.13(9)*	0.11(6)*
Ef	0.06(6)*	0.05(8)	0.04(7)
Mc	0.08(2)	0.03(9)	0.08(7)
On	0.06(4)	0.05(8)	0.11(7)*
Tr	0.07(6)	0.03(9)	0.04(7)
Wb	0.03(6)	0.02(9)	0.09(6)*
Ys		0.03(8)	0.04(7)
Nt	0.07(6)	0.10(8)	0.07(9)
Hh	0.03(3)	0.04(3)	
Ur	0.14(3)*	0.07(6)	
Sh	0.10(4)	0.05(6)	
Bd	0.07(3)	0.05(6)	
Zc	0.06(3)	0.14(5)*	
Sv			
Ar	0.05(2)		
Ro	0.05(1)		

## **Session 2/2012**

		========
18 cm	3.6 cm	6 cm
0.05(8)		Failure*
0.03(8)	0.03(3)	0.03(6)
0.12(8)	0.05(3)	0.05(6)
0.08(8)	0.04(3)	0.05(4)
0.08(7)		0.04(5)
0.03(8)	0.08(3)	0.03(6)
	0.04(3)	0.04(5)
0.08(8)	0.10(3)	0.10(6)
0.12(2)		0.04(2)
0.11(4)*	0.18(3)*	0.03(5)
0.07(4)	0.10(2)*	0.04(4)
0.09(4)	0.14(2)*	0.17(4)*
0.11(4)	0.06(2)	0.07(4)
	0.05(8) 0.03(8) 0.12(8) 0.08(8) 0.08(7) 0.03(8) 0.08(8) 0.12(2) 0.11(4)* 0.07(4) 0.09(4)	0.05(8)  0.03(8) 0.03(3) 0.12(8) 0.05(3) 0.08(8) 0.04(3) 0.08(7) 0.03(8) 0.04(3) 0.04(3) 0.04(3) 0.04(3) 0.12(2) 0.11(4) * 0.18(3) *  0.07(4) 0.09(4) 0.14(2) *

### **Session 1/2012**

Station	18 cm	3.6 cm	6 cm
ль1	6.23(2)		0.16(6)*
Ef	0.06(2)	0.05(3)	0.07(9)
Mc	0.10(2)	0.10(3)	0.08(5)
On	0.05(2)	0.10(3)	0.12(6)
Tr	0.16(2)*		0.05(6)
Wb	0.04(2)	0.06(3)	0.04(9)
Ys		0.04(2)	0.07(8)
Hh	0.05(2)	0.09(2)	0.08(8)
Ur	0.07(2)	0.10(2)	0.05(4)
Sh		0.09(2)	0.07(2)
Bd	0.05(2)	0.06(2)	0.05(6)
Sv	0.04(2)	0.10(2)	2.70(9)*
Zc	0.06(2)	0.04(2)	0.08(9)

Numbers here are <u>the median absolute error</u> in the antenna gain amplitude. This number will be approx half the error in the SEFD and is the same that you see in AIPS gain plots. The number in brackets after each entry is the number of experiments that were used to determine the median error for that entry.

### ## Effelsberg

- -- Not as good as before because DBBC 16 MHz filter caused the over-high correlation amplitude.
- -- Ways of fixing the problem in AIPS.
- (1) Run ACCOR to re-normalize the cross-correlation amplitude if the side channels were removed via FITLD.
- (2) Do not solve/apply bandpass amplitude solutions if the side channels are kept to maximise the sensitivity.

#### ## Jodrell Bank 1 and 2

- -- Jb2 sensitivity back to normal level, Tsys ~ 50 K at 18/6cm
- -- Jb2 sensitivity is ~3600 Jy at 22 GHz in Session 3/2012 No pointing calibration for Jb2.
- -- Jb1 sensitivity is good at 18cm, while not clear at 6cm
- -- No reliable rxg files available
- -- Problem with installing antabfs.pl

#### ## Westerbork at 5cm

- -- Single-dish was used at 5cm
- -- Requesting a correction factor of ~4 for Tsys in ES069 spectral line observations. Likey, there were significant sensitivity loss in the mode (2x2 MHz, RCP and LCP).

### ## Urumqi

- -- Calibration measurements should be checked carefully
- -- Low correlation amplitude in a certain BBCs.

### ## Sv, Bd, and Zc

- -- Nominal SEFD was used in the antabfs file
- -- Good calibration depends on its system stability.

#### ## Torun

- -- Good calibration was maintained since Session 2/2012
- -- Problems with receivers are less seen.

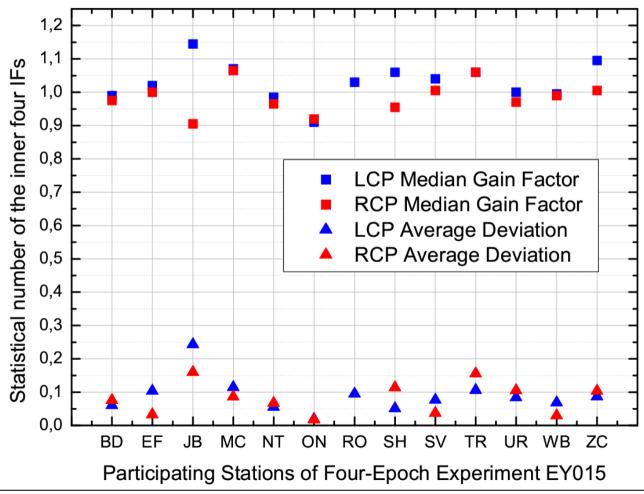
#### ## Onsala

-- Tsys was slightly (~10%) higher at 5cm in Session 3/2012

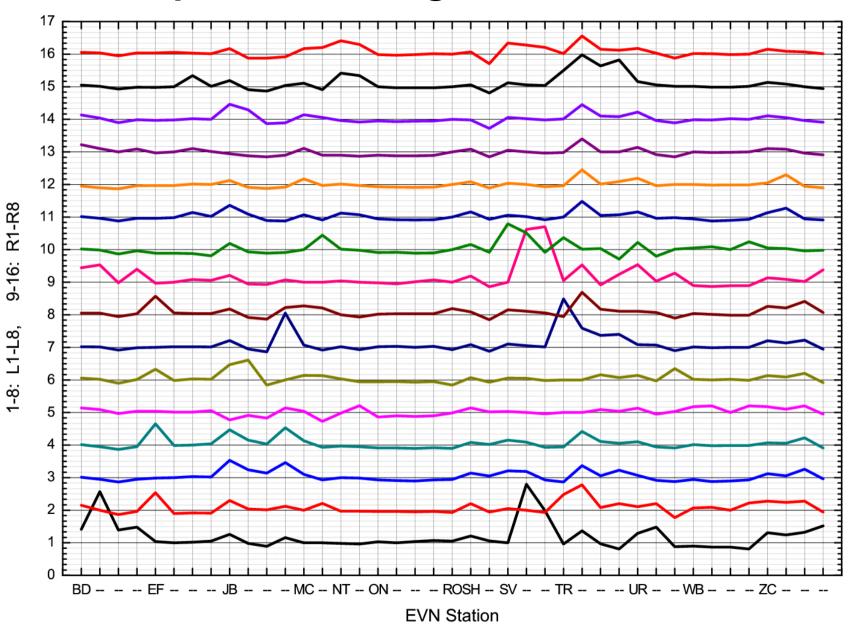
# AIPS gain factors in EY015

## 1.6 GHz, 1024 Mbps (8 subands, dual polarisations), 4 epochs. ## Determined mannually in Difmap via gscale.

## Only the inner 4 IFs were used in the statistics.



## IF-dependent AIPS gain factor in EY015



## **EVN 22 GHz status**

```
Stations | SEFD | Frequency Range | Gain Curve | DPFU | Tsys | 2Gbps
     (Jv) \mid (MHz, R\&L) \mid (Opacity-free) \mid (K/Jv) \mid
                                                  (K) | (Y/N)
        90 | 21700 - 24300 | Yes
   Ro | 83 | 18000 - 26000 | Not yet?
   Ys | 200 | 18000 - 26000 | Yes
                                                 50 | Yes
   On | 1380 | 19000 - 26000 | Yes
                                                | 150 | Yes
   Nt | 800 | 21950 - 22325 | Not vet?
   Hh | 3000 | 21400 - 24100 | No
                                         10.07 | 300 | Yes
   Mh | 2600 | 21980 - 22480 | No
                                         10.03
                                               | 100 | Yes
   Tr | 500 | 22000 - 22500 | Not yet?
   Mc | 900? | 18000 - 26000 | Not yet?
           | 22180 - 22280 | No Tsys data |
    Ur | 850 | 22100 - 24000 | Not vet?
                                          10.07
    Sh | 1700 | 22100 - 22600 | Yes
        1143 | 22020 – 22520 | No Tsys data
              22020 - 22520 | No Tsys data |
              2020 - 22520 | No Tsvs data
                                           0.08
              21250 - 23250 | Yes
              21250 - 23250 | Yes
                                          0.08
                                                       5B+
              21250 – 23250 l
                                           0.08
                                                        5B+
         934 |
                               Yes
```

#### Note

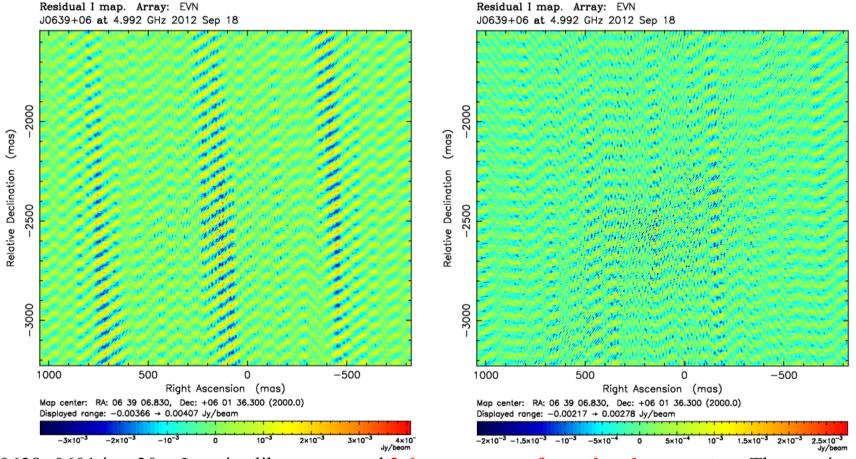
- (1) Only RCP available at Noto.
- (2) Tsys data includes opacity correction at Onsala via the chopper wheel calibration.

## New antabfs program

## Supporting DBBC on-off Tsys measurement and 80Hz radiometry

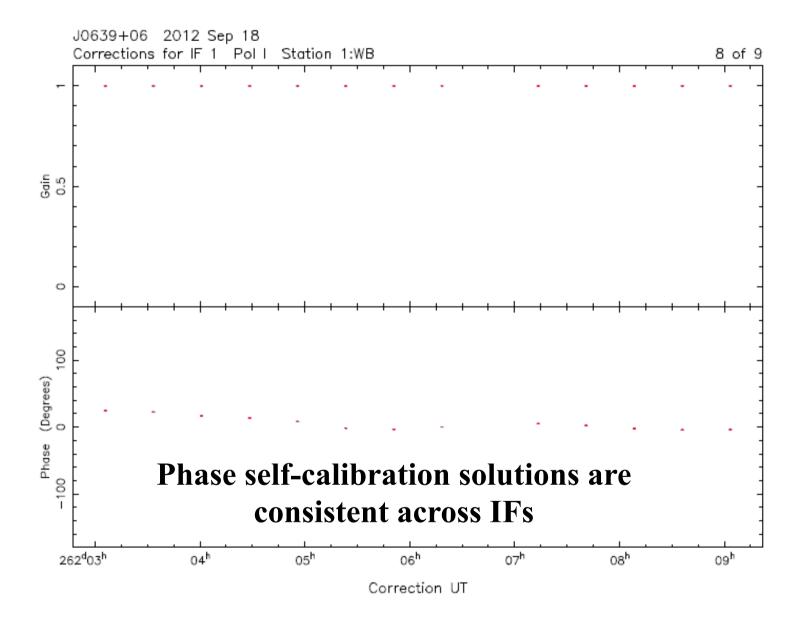
- ## Supporting mix-mode wide-band VLBI observations
- -- Parsing vex file as well to get the default frequency setup information
- ## Adding scan-dependent smoothing or curve fitting to reduce the noise.
- -- Requesting data editing as less as possible
- ## Trying to use wide-band Tsys measurement and Tsys-frequency curve to derive Tsys for each subband
- -- Less affected by RFI
- -- More sensitive for small telescopes
- ## Adding an input file for the EVN pipeline to easily correct the residual amplitude errors of each subband
- -- Very useful for the rapid e-VLBI observations

# Phase error related to WSRT digital beam?



## J0639+0601 is a 30 mJy point-like source and **2.6 arcsec away from the phase center**. These stripes in the left residual map can be removed by the phase self-calibration of Wb data. See the right panel.

## These stripes are not seen any more in the residual map of phase-referencing observations in the later epochs when the source position error was corrected.



# Timely delivery of data

Timely delivery can significantly speed up the correlation, post review, and pipeline processes and make more disk packs be available in the upcoming session.

Feedbacks, rxg, and antab files should be delivered within 2 weeks after the end of a session and ASAP in the case of e-VLBI experiments.

Automatically uploading log files and gps data are very welcome.