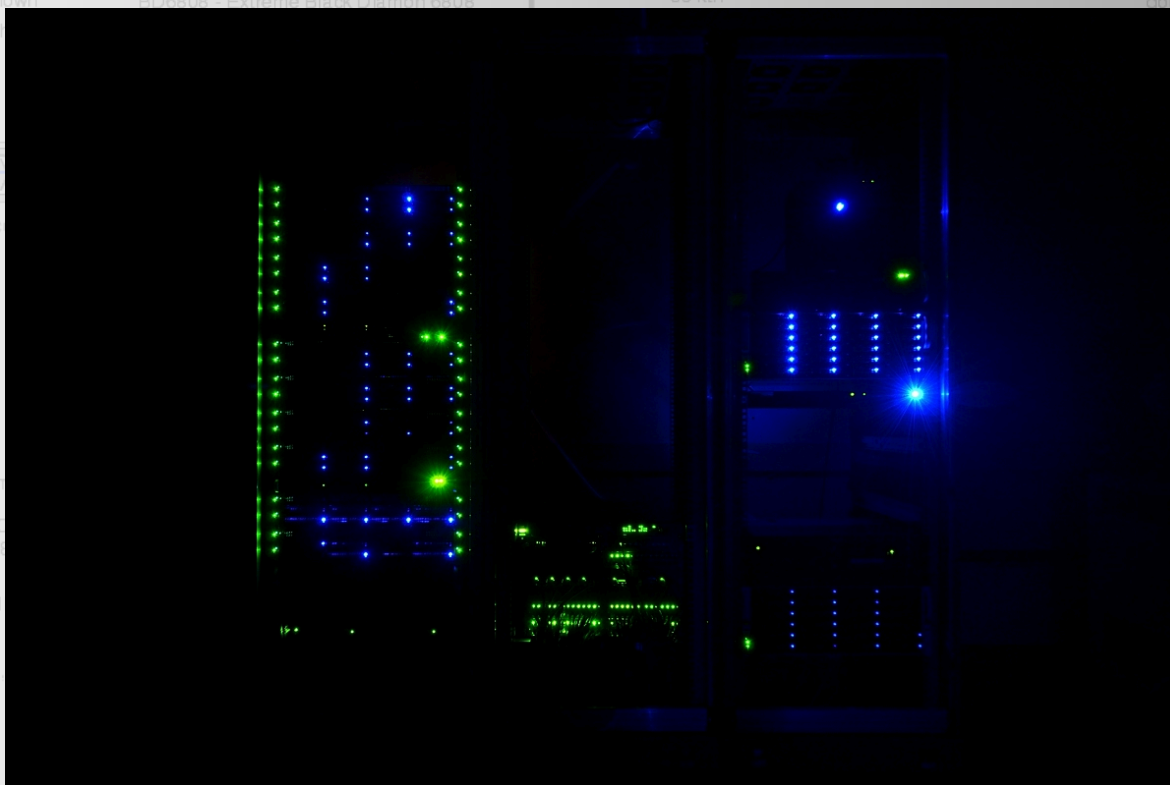
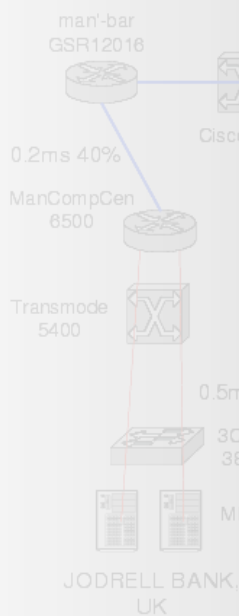


Routes across GEANT
used by eVLBI MkVs

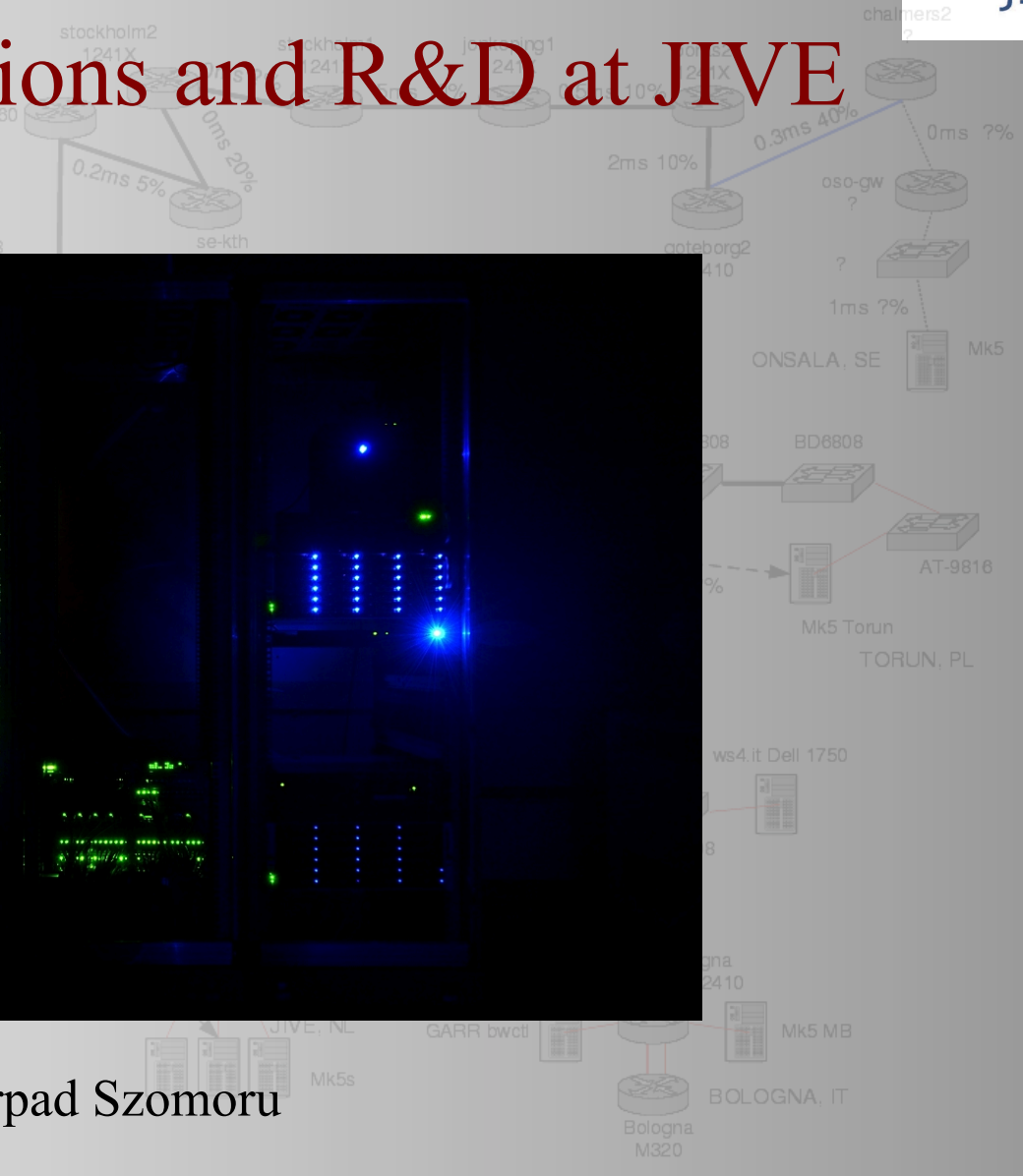
Technical Operations and R&D at JIVE

RTT & % load	Equipment
—	STM-64 (10Gbps)
—	STM-16 (2.5Gbps)
—	Gigabit Ethernet
.....	Unknown

% load is approx daily high



Arpad Szomoru



What do we do?



- **Correlators**
 - More capacity, new telescopes, development of AVN
 - New features, new science
- **Data recording/playback/transport**
 - Real time/near-real time
 - Higher bandwidths, 2 and 4Gbps
- **Automated operations**
 - Get rid of disk shipping
 - Monitoring, automated fringe checking
 - Triggered observations
- **SKA and mm VLBI**
 - **User software**, VLBI with CASA
 - Simulations for BHC
 - Fringe checking
- **Time and frequency transfer**
 - For SKA
 - And on public networks

Correlation

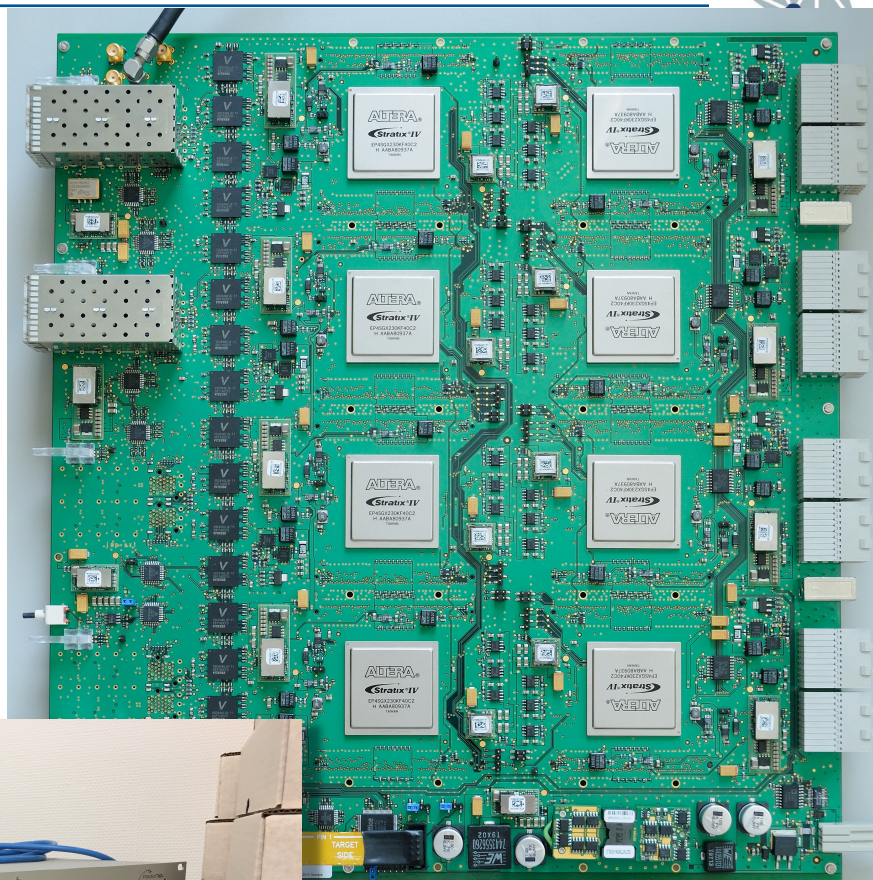
- SFXC software correlator at JIVE:
- 43 nodes; 452 cores
(Intel Xeon 5500/5600/E5-2600/E5-2630)
- QDR Infiniband interconnect
(40 Mbit/s)
- 8 nodes with 10 GbE
(currently limited to 30 Gbit/s total)
- 15-16 stations @1Gbit/s real-time
(with cross-polarisations)
- All recorded VLBI on SFXC since summer 2012
- First real-time e-VLBI in december 2012
- New SFXC features: talk by Mark Kettenis



JUC (JIVE UniBoard Correlator)



- Developed in RadioNet FP7 (also in RadioNet3)
- As of last week correlator completely integrated in control system (operator-proof) (?)
- 32 MHz sub-bands: last bugs
- Start correlating from Mark5 recordings (again)
- Real-time tests



16 stations at 4 Gbps?



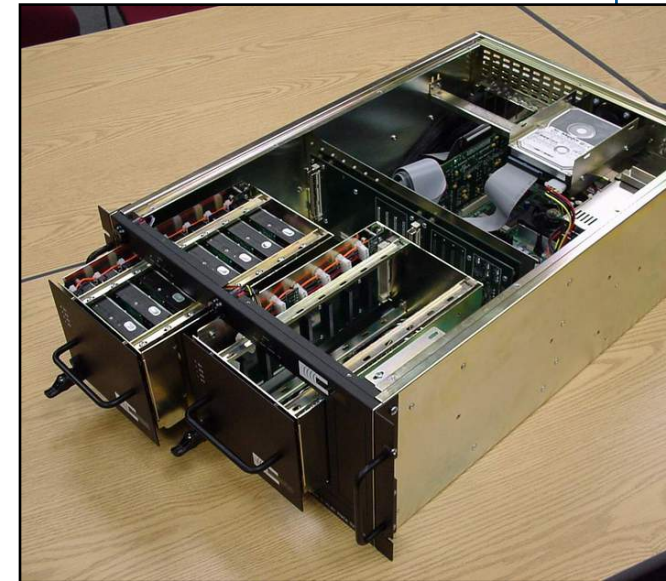
- ~13 stations on current SFXC cluster at 1 Gbps
- 16 stations at 4 Gbps: factor of ~5 more hardware needed

- 16 stations on 2 available UniBoards at 2 Gbps
- 4 Gbps: factor 2

platform	power consumption (kW)	investment (keuro)
SFXC	30	550
UniBoard	1	30

- Somewhat incomplete comparison
 - Software: fantastically flexible, easy to modify, HW getting cheaper as we go along
 - Firmware: very power efficient, once it goes, it goes, but not nearly as flexible, ideal for “simple” things

- Original control code not adequate for needs of EVN
 - Full re-write of Mark5 control code
 - Used for all EVN operations, gaining traction in geo community as well
 - Incorporates full Mark5 command set, supports Mark5, Mark6, FlexBuff...
 - Many features; “Swiss army knife of (e)VLBI”
 - Channel dropping, on-the-fly corner turning, sending different chunks of data to different destinations, full VDIF support
 - Made e-VLBI possible at all
 - Enabled semi-automated fringe tests
 - M5Copy: transport any data from anything to anything
 - Choice between TCP, UDP, UDT
 - Essential for automated shipping
 - Future developments in ASTERICS
 - Talk by Harro Verkouter



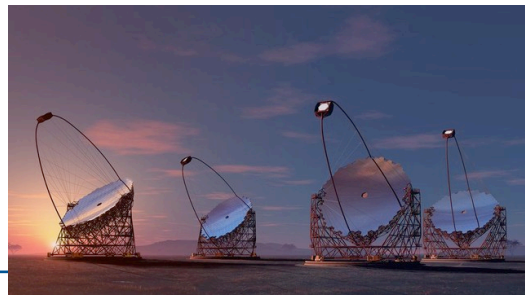
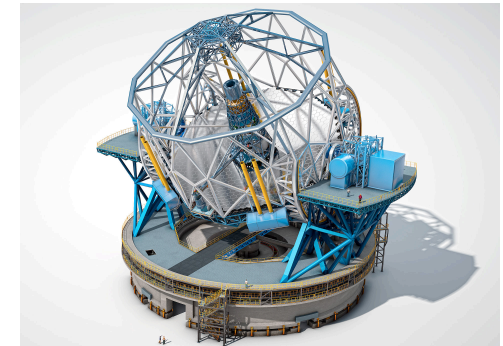
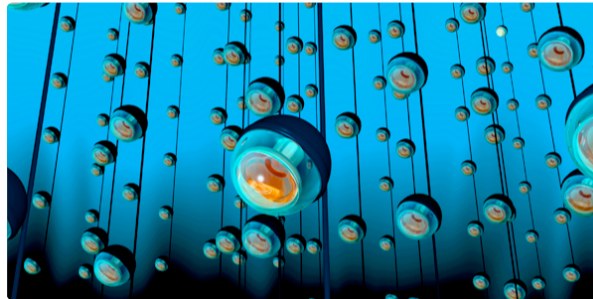
Asterics

Astronomy ESFRI & Research Infrastructure Cluster

Addressing Cross-Cutting Synergies and Common Challenges for
the Next Decade Astronomy Facilities

What is ASTERICS ?

- Astronomy ESFRI Research Infrastructure CluSter (ASTERICS)
- Topic: Implementation of cross-cutting solutions for clusters of ESFRI research infrastructures and ERICs
- ASTERICS represents the first major European collaboration Astronomy/Astrophysics/Astroparticle Physics
- **23 partners funded by EC Horizon2020 at 15 M€ for 4 years**



WP2 – DECS: Dissemination, Engagement and Citizen Science

WP3 - OBELICS: Observatory E-environments Linked by Common challenges

- Software interoperability
- Training in parallel programming and big data frameworks
- Adapt and optimise extremely large database systems for ESFRIs
- Data mining tools and statistical analysis techniques on petabyte data sets

WP4 - DADI: Data Access, Discovery and Interoperability

- Training and support for the scientific use of VO in general
- Train and support staff of ESFRI projects
- Adapt VO framework and tools to ESFRI projects needs

WP5 - CLEOPATRA: Connecting Locations of ESFRI Observatories and Partners in Astronomy for Timing and Real-time Alerts

- Time and frequency transfer
- relaying alerts (warning system for transient events, also in EVN)
- data streaming software (builds on Jive5ab experience)
- advanced scheduling algorithms for complex, large arrays (mainly for SKA, CTA)

Time and frequency transport: White Rabbit



- CLEOPATRA:
 - Verify/demonstrate achieved 10^{-13} stability (1s) and 1ns timing performance
 - by showing fringes between WSRT and Dwingeloo dish
 - Transfer of H-Maser signal from the WSRT to Dwingeloo
- But also used for SKA:
 - Measure timing performance with WR-ZEN board
 - Test with 10km fibre in climate chamber
 - Tests done on overhead fibres in SA
 - Tests planned on e-Merlin fibre



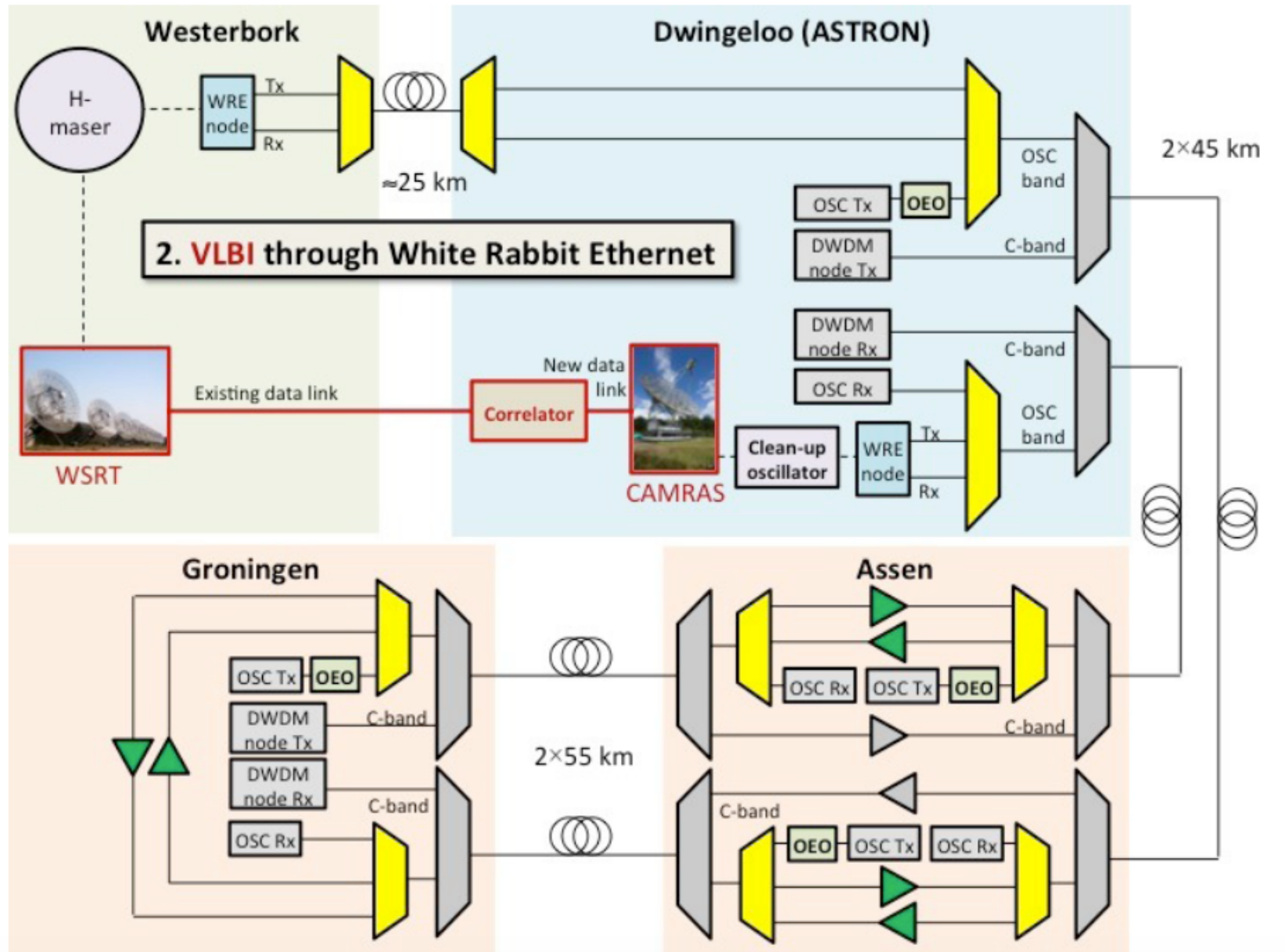
What is White Rabbit?



- Sub-ns accurate synchronization network
- Open Hardware design, project started at CERN
- Based on:
 - PTP (IEEE1588v2)
 - Bidirectional (BiDi) SFPs
 - SyncE: Syntonization of 125 MHz clock
 - 1 Gb/s Ethernet
- In use in several accelerators and astronomy instruments around the world



Frequency transfer demo in CLEOPATRA

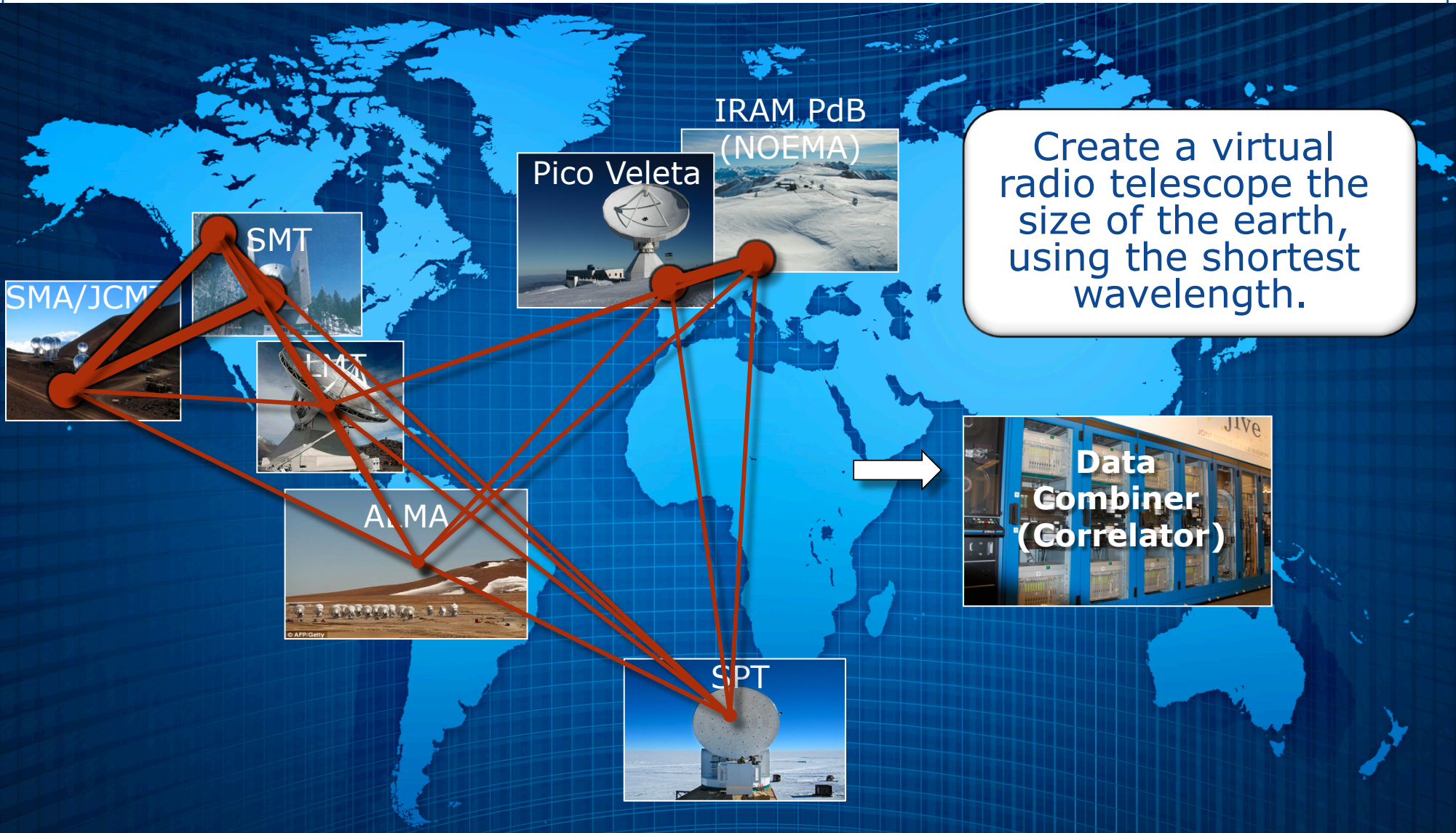




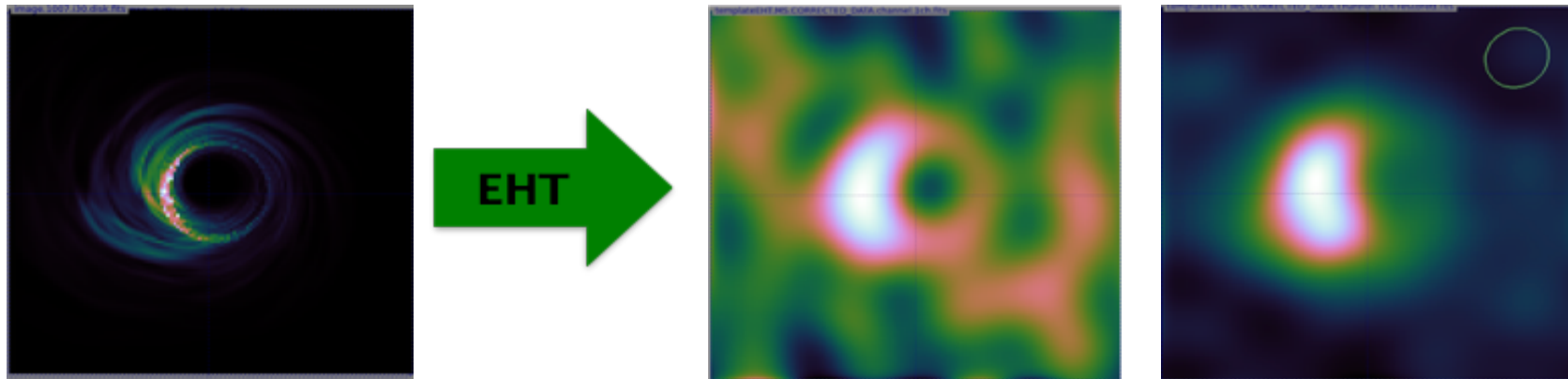
Desolation and overhead fibres



BlackHoleCam and the Event Horizon Telescope



- Software pipeline [JIVE]
 - CASA-based fringe-fitter for VLBI
 - End-to-end VLBI reduction in CASA now possible (thanks to many hacks)
- Array simulation [JIVE, Radboud, Rhodes]
- Robust turn-key VLBI operation [MPIfR, JIVE]
 - “real-time-ish” fringe check mechanism in place (based on jive5ab)
 - Interface being implemented, gets info from observing schedule
 - Uses dead time between scans, continues data transfer over many intervals
 - No interference with recording in any way



Courtesy: Monika Moscibrodzka, Roger Deane

mm-VLBI fringe checking



- Based on experiment schedule
- Select scans, gaps for transfer

The screenshot displays the 'Fringe test scheduler' application. It features a 'Choose VEX file:' field with a 'Browse' button. Below this is a 'Stations' section with a table of scan data. A file selection dialog is open, showing the file 'c152b.vex' selected in the 'Files of type: VEX' list. The main table contains the following data:

Scan	Start time	End time	Source	Stations	Mode
No0500	2015y267d23h56m58s	2015y267d23h57m58.000000s	P-CHICYG	...Fd..KpLa...NL..Pt..	PulseCalOFF
No0502	2015y267d23h57m22s	2015y267d23h58m22.000000s	SIO-RCAS	Br.....Ov.....	PulseCalOFF#02
No0501	2015y267d23h58m04s	2015y267d23h59m04.000000s	P-CHICYG	...Fd..KpLa...NL..Pt..	PulseCalOFF
No0503	2015y267d23h58m28s	2015y267d23h59m28.000000s	P-CHICYG	...Fd..KpLa...NL..Pt..	PulseCalOFF
No0504	2015y268d00h00m00s	2015y268d00h00m00.000000s	BLLAC	aMh..NL0vPtYs	3mm_RDBE
No0505	2015y268d00h07m30s	2015y268d00h07m30.000000s	BLLAC	...Gb.....	3mm_RDBE
No0506	2015y268d00h12m14s	2015y268d00h12m14.000000s	SIO-RCAS	...Fd..KpLa...NL..Pt..	PulseCalOFF#03
No0508	2015y268d00h12m21s	2015y268d00h12m21.000000s	SIO-RCAS	Br.....Ov.....	PulseCalOFF#02
No0507	2015y268d00h13m20s	2015y268d00h13m20.000000s	SIO-RCAS	...Fd..KpLa...NL..Pt..	PulseCalOFF#03
No0509	2015y268d00h13m27s	2015y268d00h13m27.000000s	SIO-RCAS	Br.....Ov.....	PulseCalOFF#02
No0510	2015y268d00h15m00s	2015y268d00h15m00.000000s	BLLAC	BrEbFdGbKpLaMh..NL0vPtYs	3mm_RDBE

The interface also includes a 'Schedule selected scan' button and an 'Allow transfers in all gaps' checkbox. A 'Schedule scan' dialog is also visible, showing a start time of '2015-09-25 00:03:30' and a duration of '4.00' seconds.

- Working on an H2020 proposal
 - Aimed specifically at ESFRI, ERICs
- **BUMPING UP JIVE:**
 - Bringing Users More Power - Interferometry's Next Global Usage Perspective by JIVE
- Objectives:
 - Solidify JIVE's presence in Europe's radio astronomy community
 - Raising the public awareness of JIVE and the EVN
 - Expand the scope of the ERIC (ILT?)
 - Further integration of new telescopes
 - Develop geodetic and astrometric capabilities of JIVE/EVN
 - Aim for a global VLBI facility
 - Operational interface issues: remote monitoring/control, SCHED?
 - large-scale exchange programme for researchers and technicians EVN-AVN