

Cosmic Microwave Background Polarization Receivers: QUIJOTE experiment

E. Artal⁽¹⁾, L. de la Fuente⁽¹⁾, B. Aja⁽¹⁾, J.L. Cano⁽¹⁾, E. Villa⁽¹⁾
R. Hoyland⁽²⁾, J.A. Rubiño⁽²⁾, R. Génova⁽²⁾

(1) Universidad de Cantabria, Santander (Spain)

(2) Instituto de Astrofísica de Canarias, La Laguna, Tenerife (Spain)

Multi Pixel Camera Receivers,
RadioNet- FP7, 2nd Engineering Forum Workshop (Bonn, 16-17 November 2009)



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QUIJOTE CMB experiment overview

- Q-U-I JOint TEnerife (Stokes parameters Q, U and I)
- Cosmic Microwave Background (CMB) polarization receivers
- To obtain five polarization maps in the frequency range 11- 30 GHz
- Angular resolution: ~1 degree

QUIJOTE experiment consortium



Instituto de Astrofísica de Canarias (IAC), Tenerife (Spain): Coordinator



Instituto de Física de Cantabria (IFCA), Santander (Spain)



Universidad de Cantabria (UC), Santander (Spain)



University of Cambridge, (UK)



University of Manchester, Jodrell Bank Centre for Astrophysics (UK)



IDOM, Bilbao (Spain)



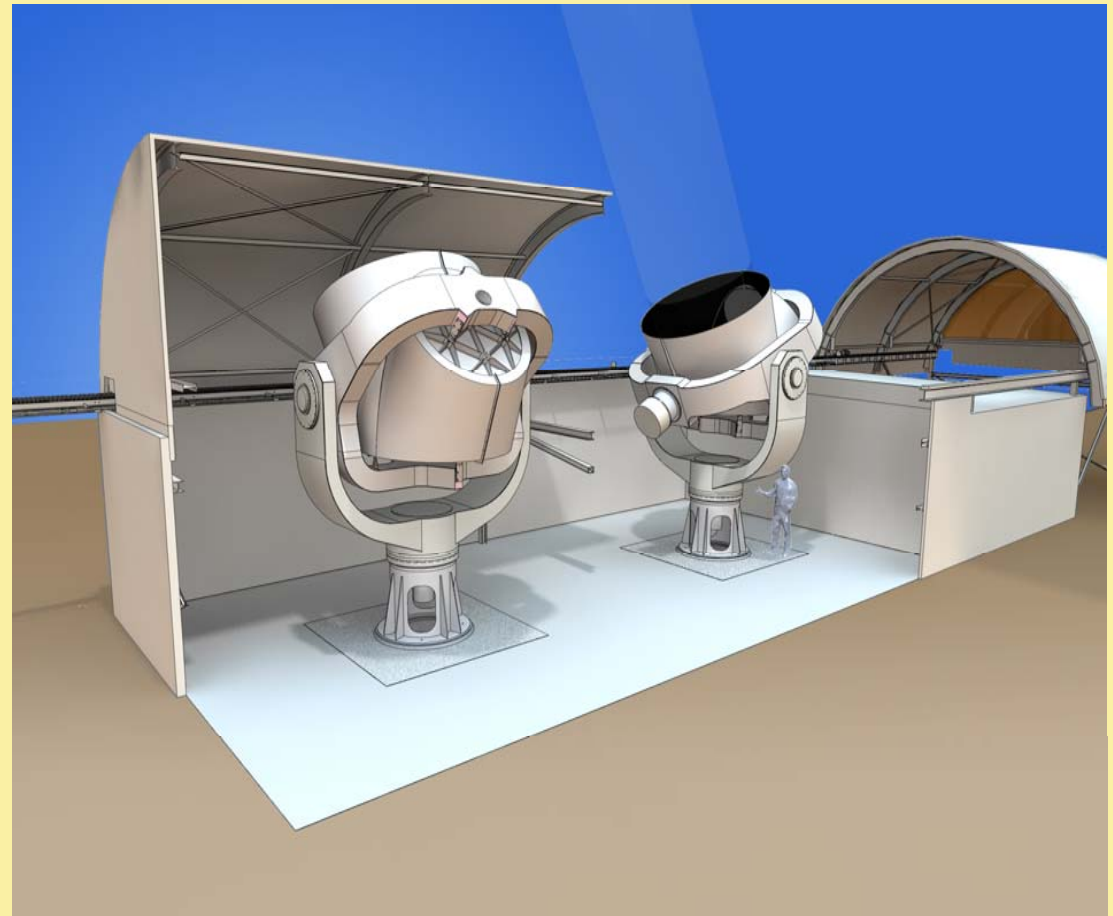
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Observatorio del Teide (Tenerife, Canary Islands)



Izaña site, 2.390 m



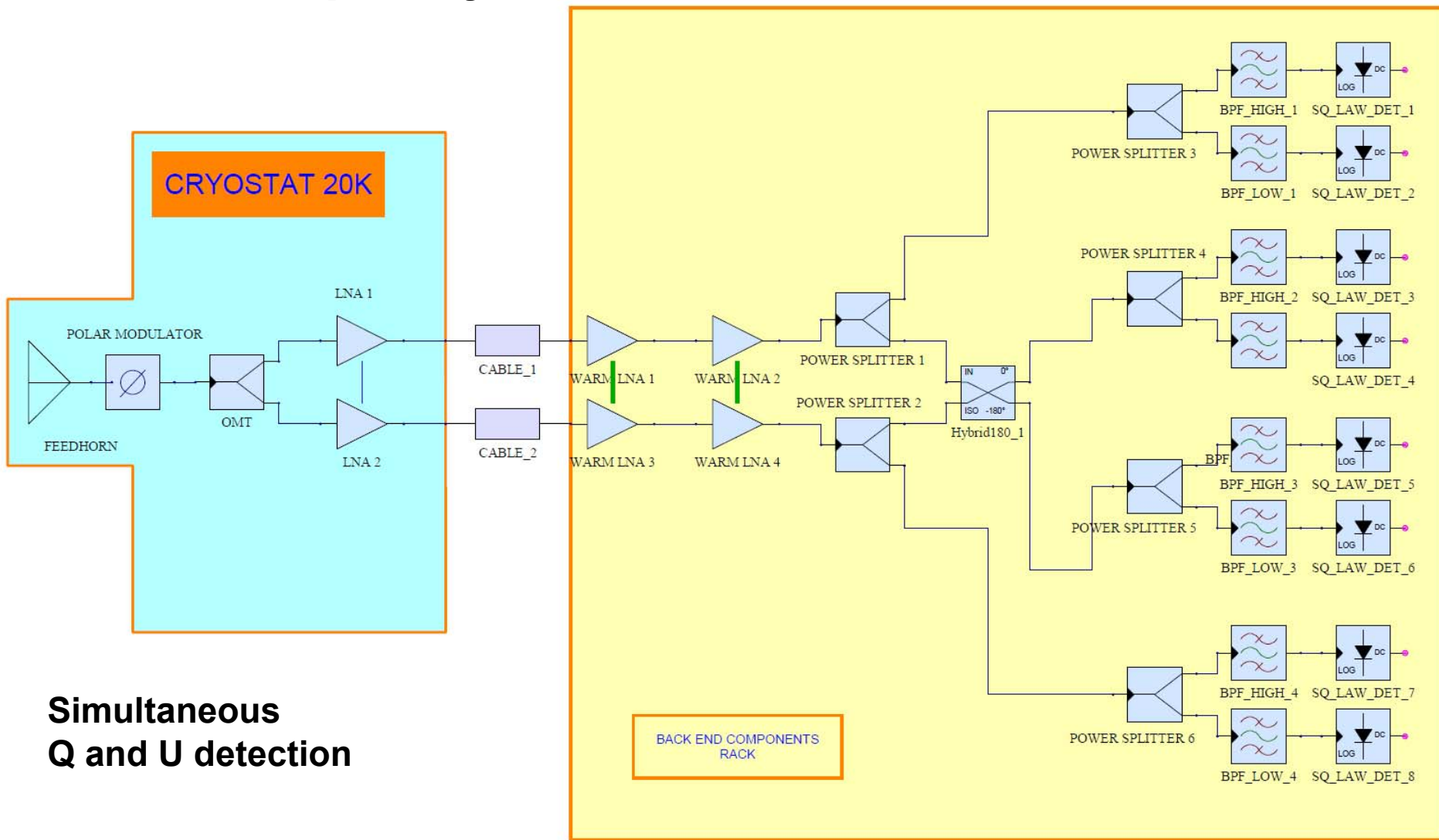
QUIJOTE Instruments 1 and 2
and enclosure

QUIJOTE Experiment. Basic features

	Instrument 1					Instrument 2
Frequency (GHz)	11.0	13.0	17.0	19.0	30.0	30.0
Bandwidth (GHz)	2.0	2.0	2.0	2.0	8.0	8.0
Number of channels	8	8	8	8	2	32
Beam FWHM (deg) (*)	0.92	0.92	0.60	0.60	0.37	0.37
T_{sys} (K)	20.0	20.0	20.0	20.0	30.0	20.0
Sensitivity ($\text{mK s}^{1/2}$)	0.22	0.22	0.22	0.22	0.34	0.05
Sensitivity per beam ($\text{Jy s}^{1/2}$)	0.24	0.34	0.24	0.30	0.43	0.07

(*) Pixel = a square with each side is FWHM (Full Width at Half Maximum) of the beam.

Low frequency channels in QUIJOTE 1



Low frequency channels: 11-13 GHz and 17-19 GHz: eight channels per pixel

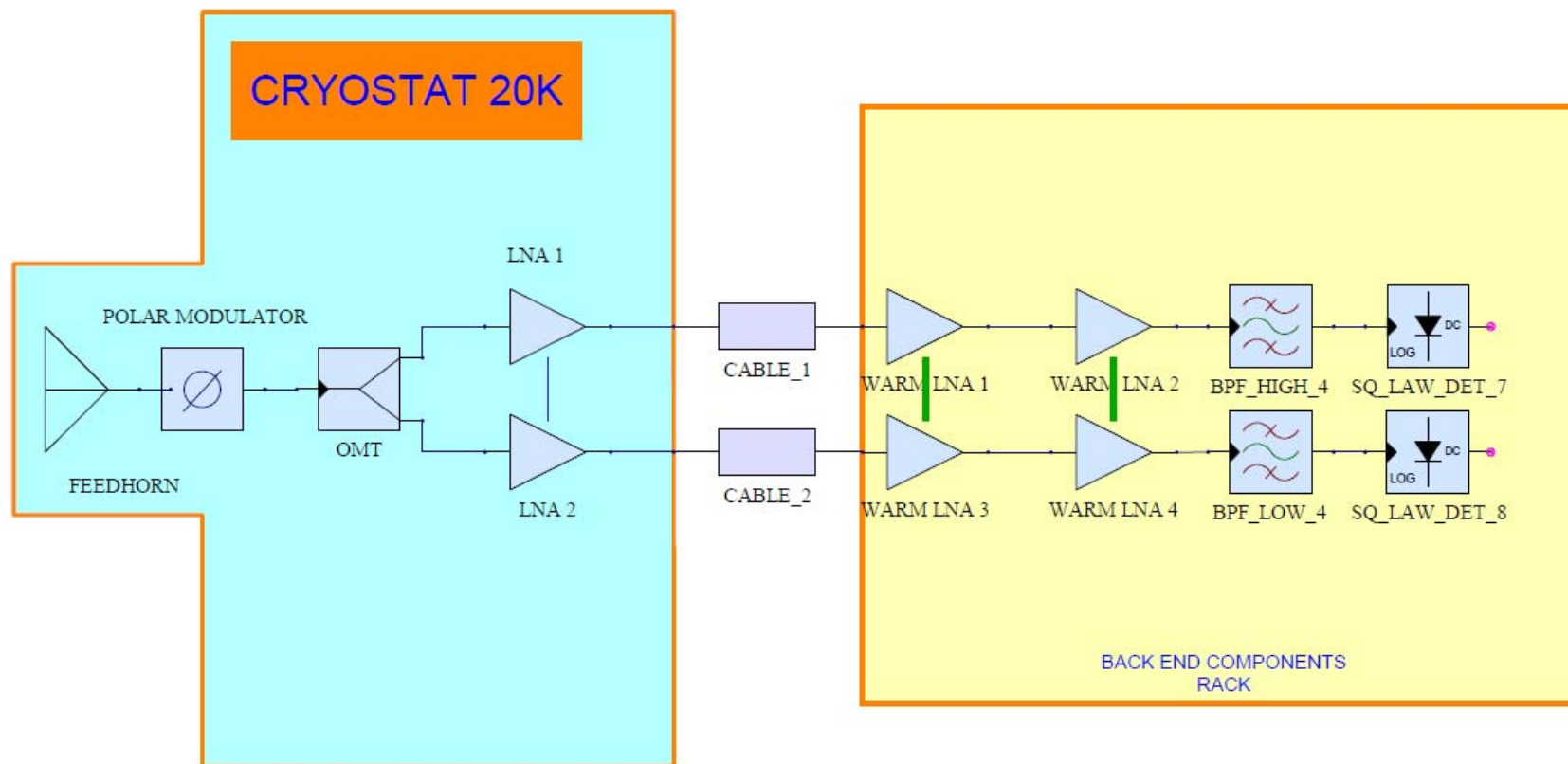
Output detected voltage in QUIJOTE 1 - Low frequency channels

$$V = \frac{1}{2} I + \frac{1}{2} Q \cos(4 \varphi) + \frac{1}{2} U \sin(4 \varphi)$$

Q, U and I = Stokes parameters
 φ = Position angle of the modulator

These parameters depend on the time
and on the channel

30 GHz channel in QUIJOTE 1

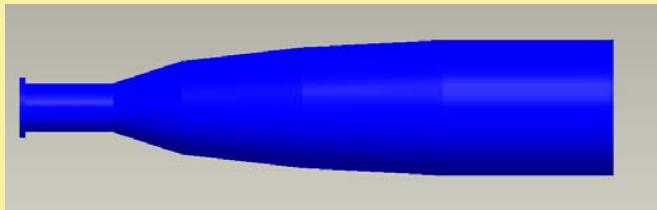


Modulator to stabilize gain drift. Simple design: two channels per pixel.

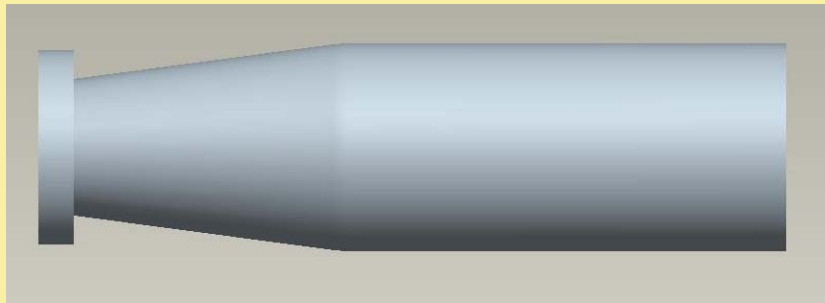
QUIJOTE 1: Focal plane distribution



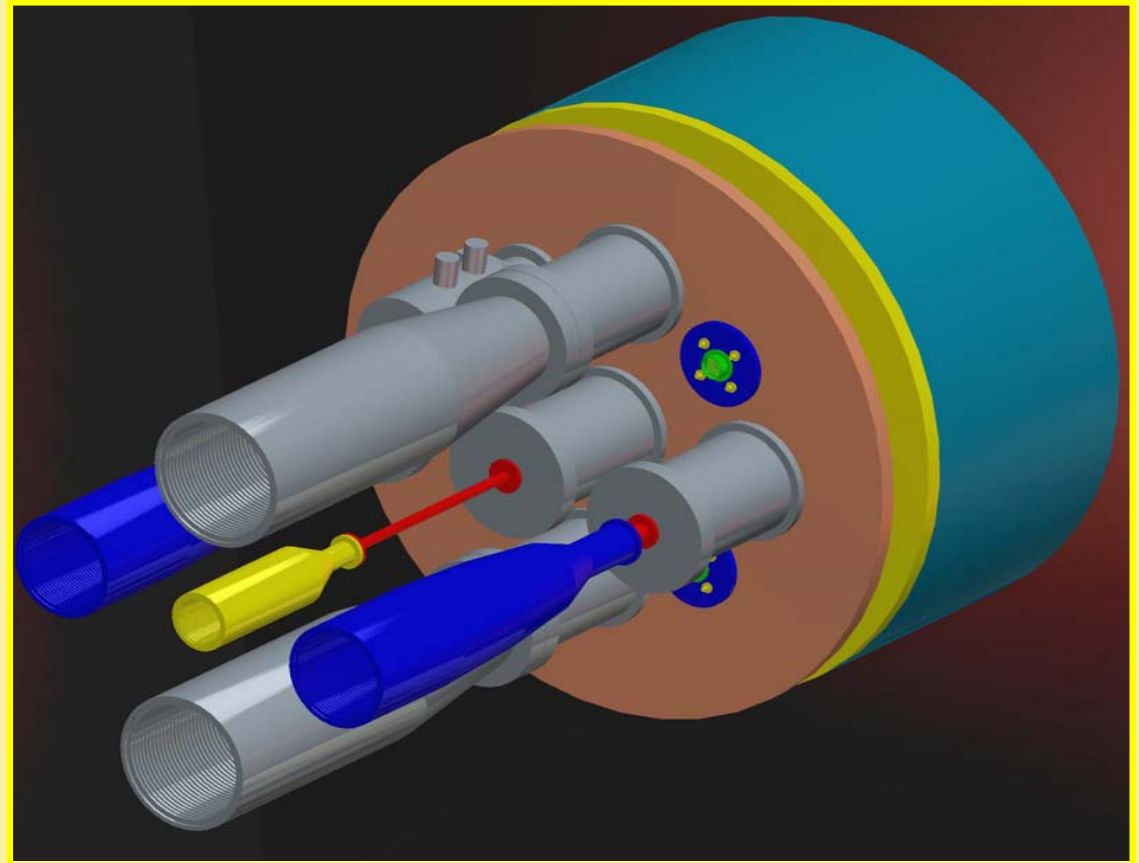
26-36GHz Horn



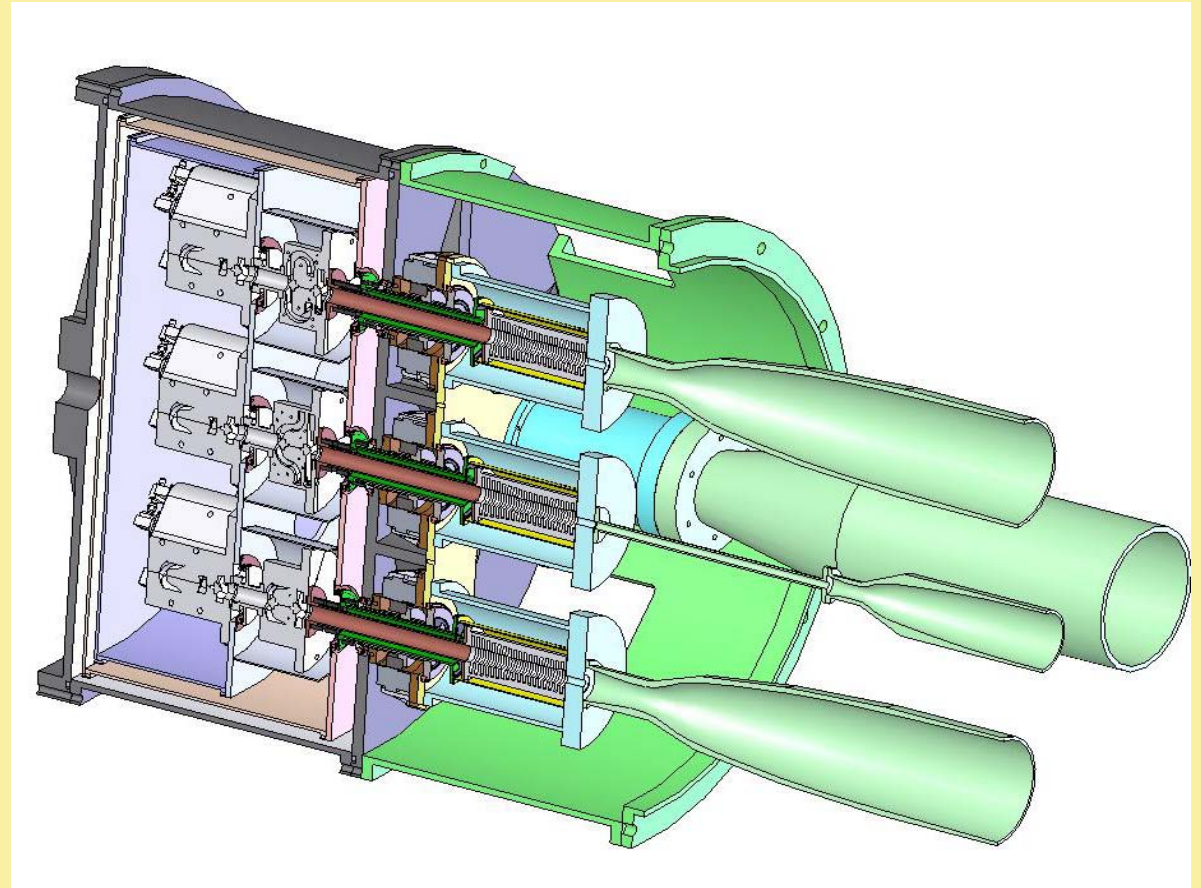
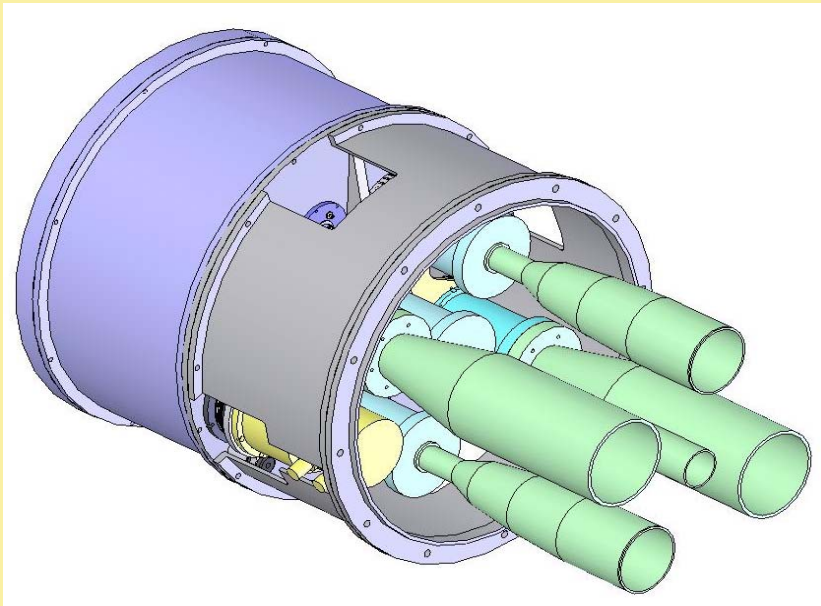
14-20GHz Horn



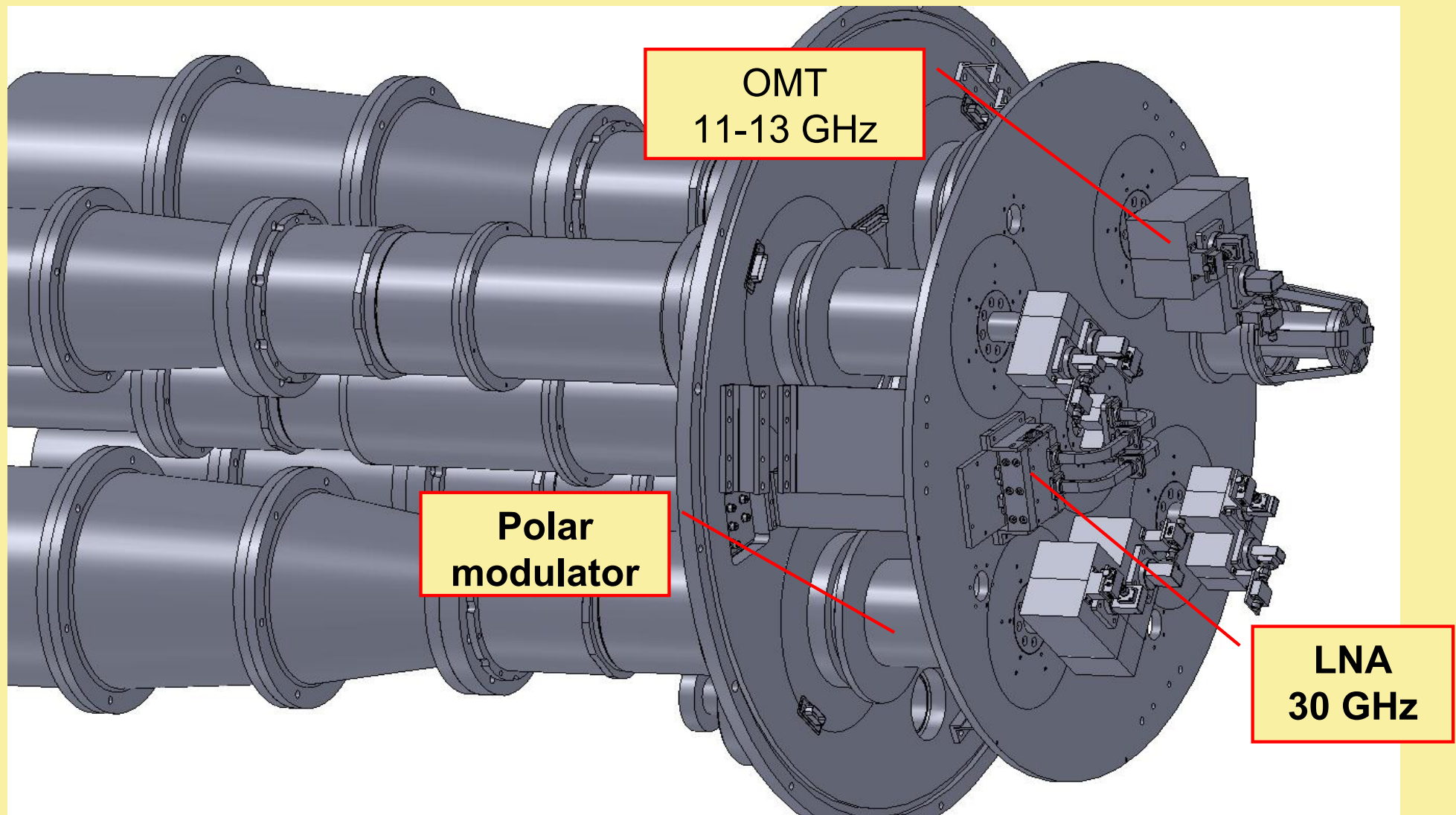
10-14GHz Horn



QUIJOTE 1: Mechanical design (version May 2008)



QUIJOTE 1: Mechanical design (version Feb. 2009)



QUIJOTE 1: Horns



Corrugated feed-horns

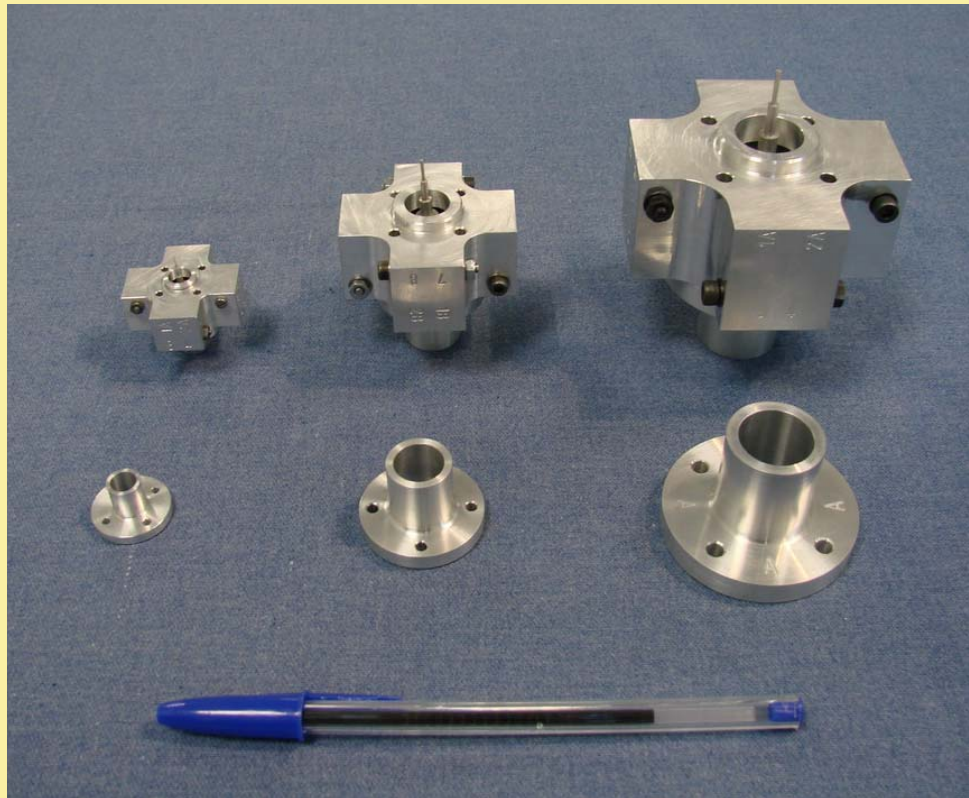
10-14 GHz

14-20 GHz

26-36 GHz

Polar Modulator

- Key component of the polarimeter
- Rotating polar modulator (40 Hz): switch out 1/f noise
- Incoming signal modulated at 4 x (modulator frequency)
- Cryogenically cooled: low losses, low impact on noise
- Waveguide component: turnstile 4-way junction



Units:

10-14 GHz

14-20 GHz

26-36 GHz

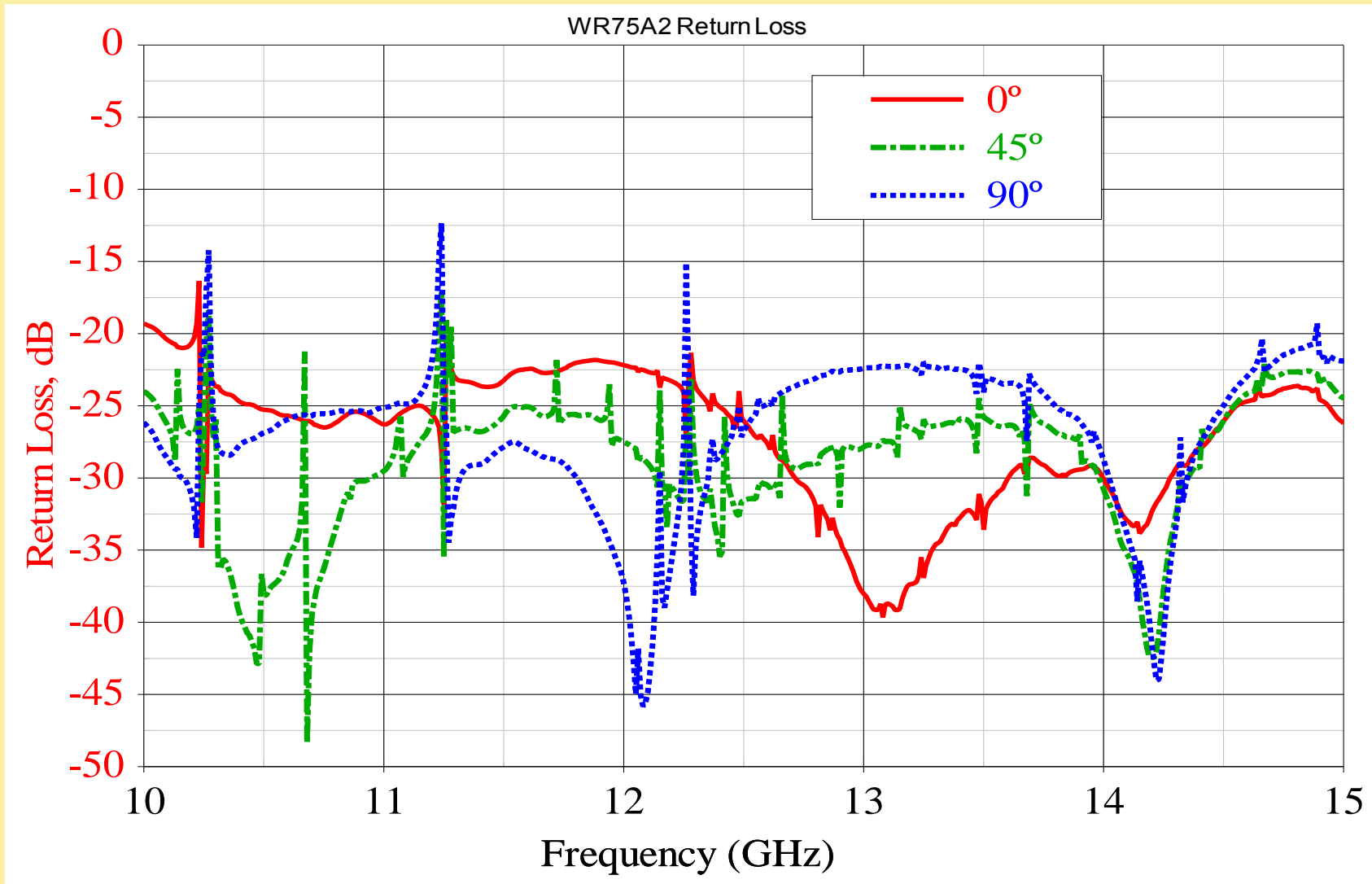


Polar Modulator S-parameter tests

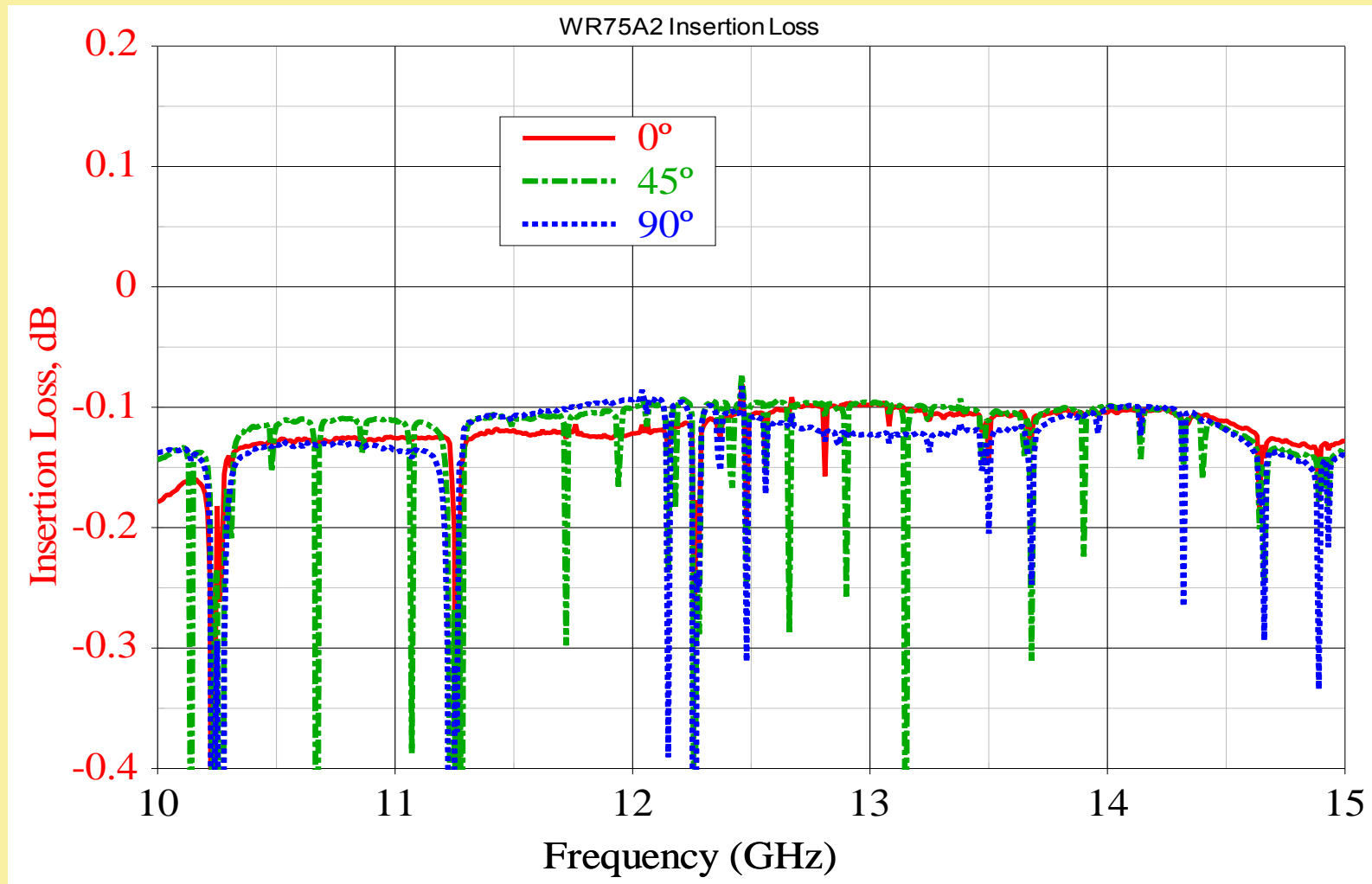


- Low return losses
- Low insertion losses
- High polar isolation
- Tests at 3 polar modulator orientations: 0° , 45° and 90°
- Plots: spikes caused by misalignment of the circular waveguide interfaces
- Tests at Room Temperature

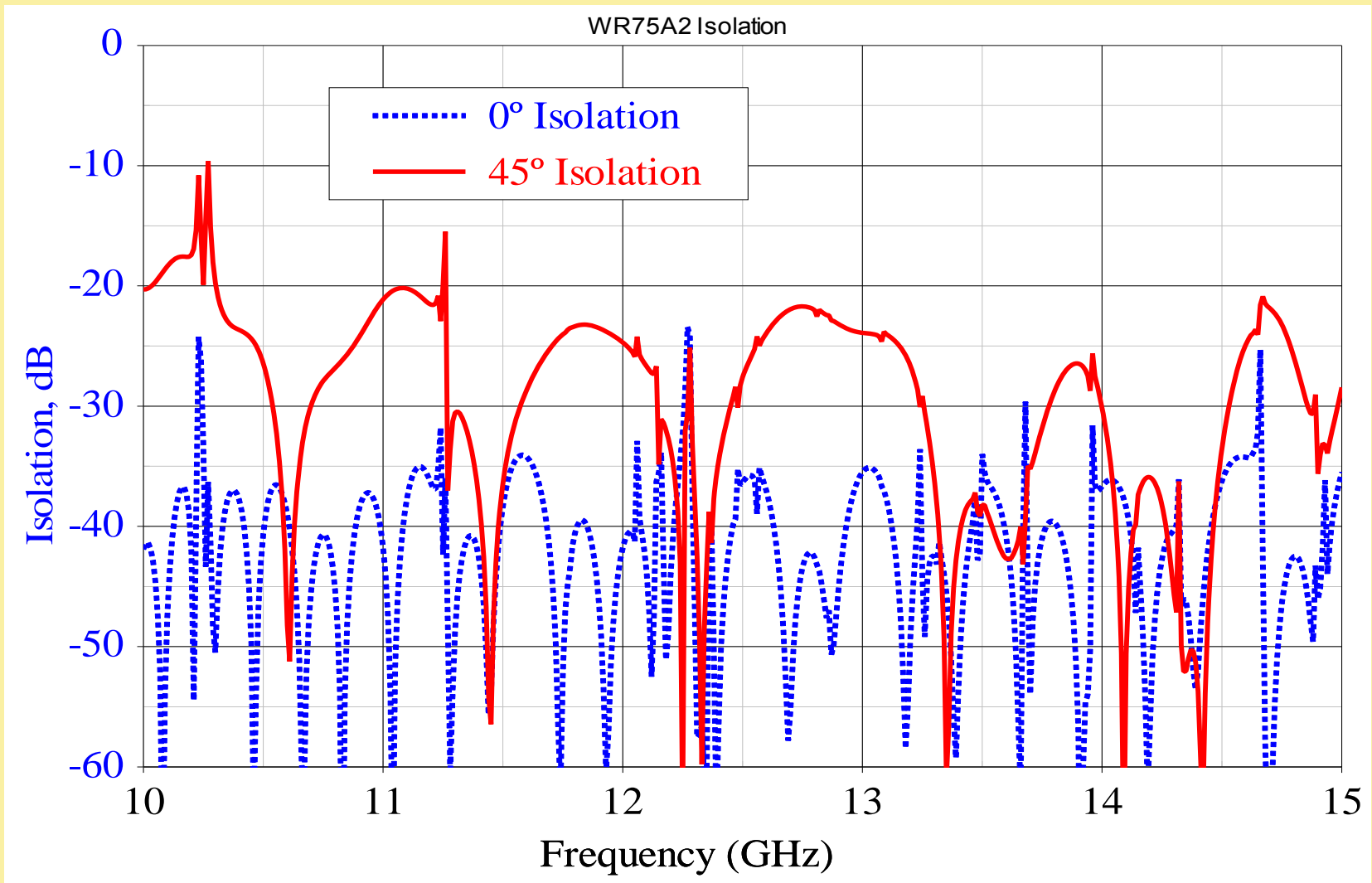
Polar Modulator Return Loss



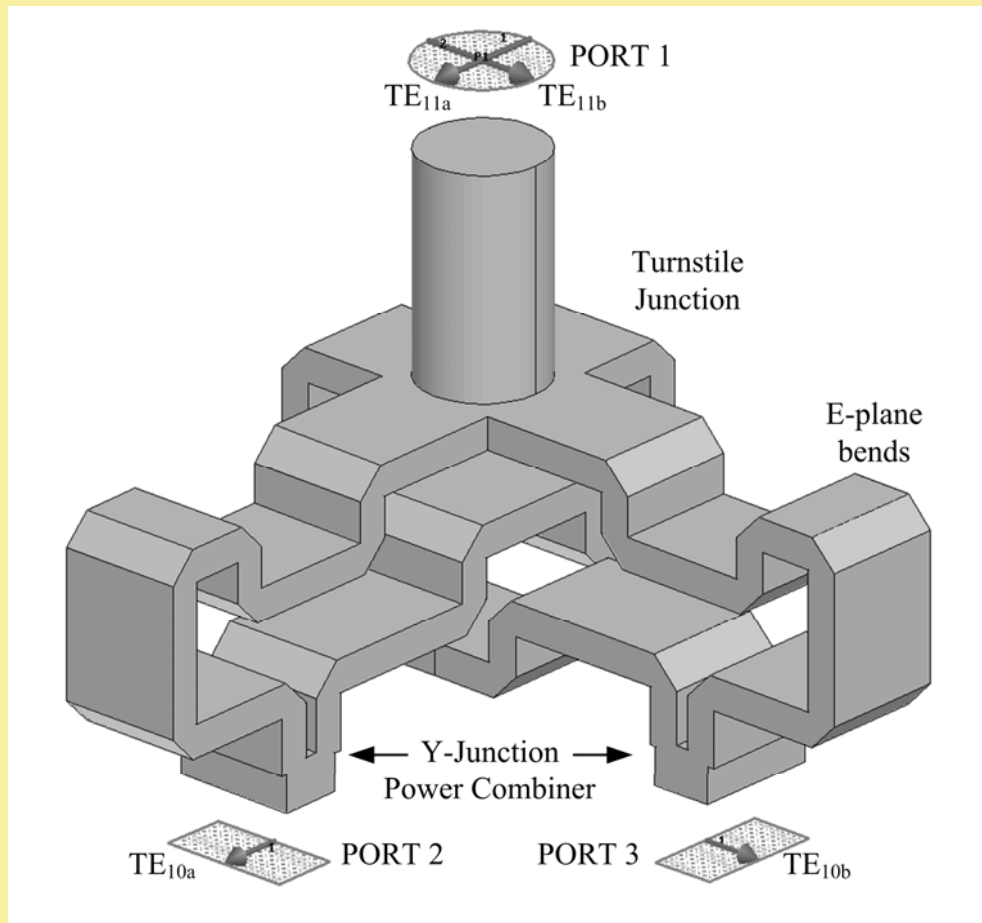
Polar Modulator Insertion Loss



Polar Modulator Isolation

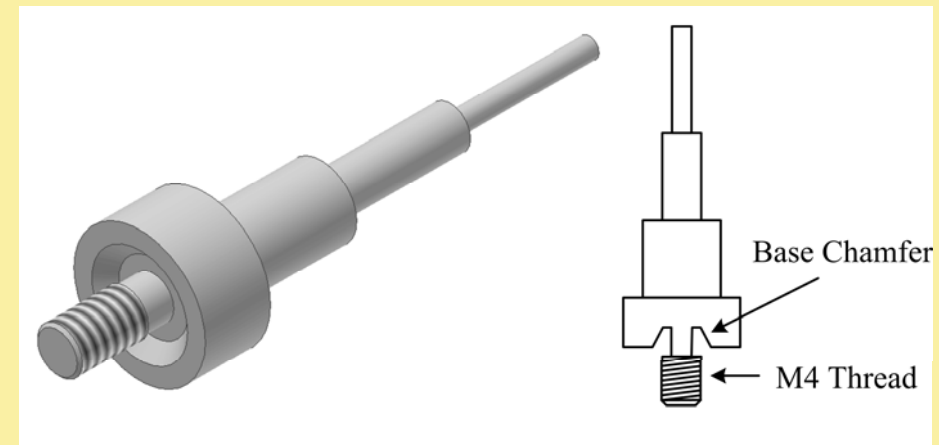
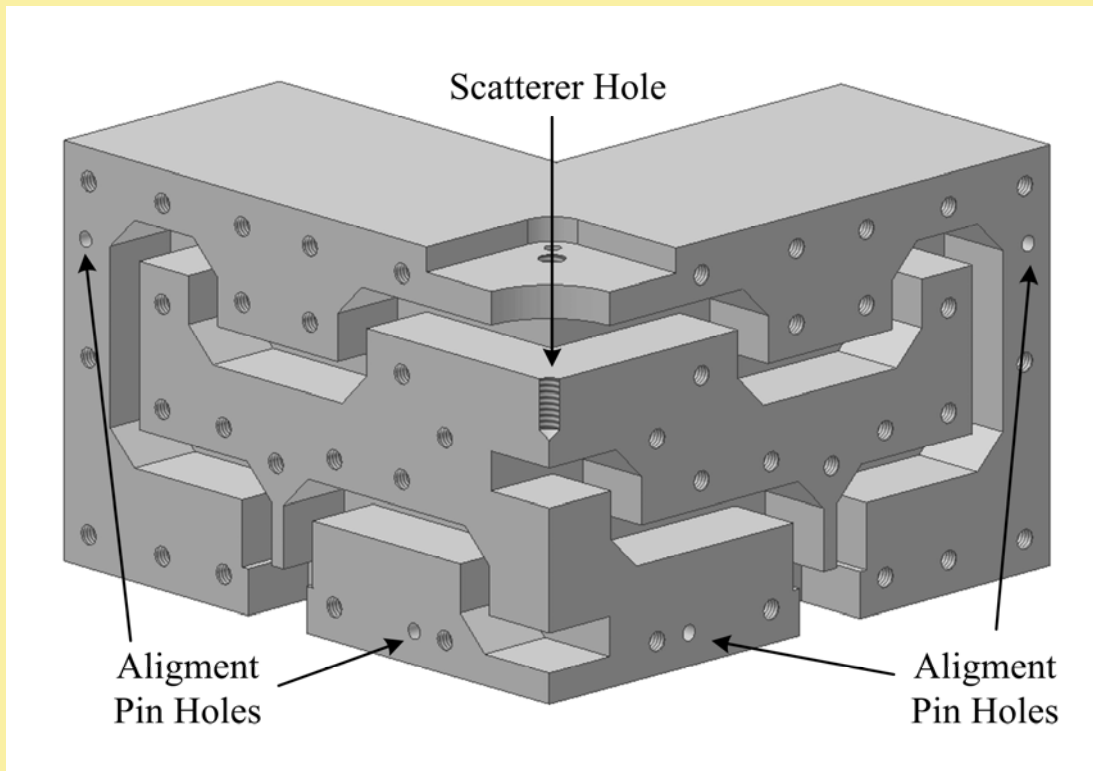


Orthomode Transducer (OMT)



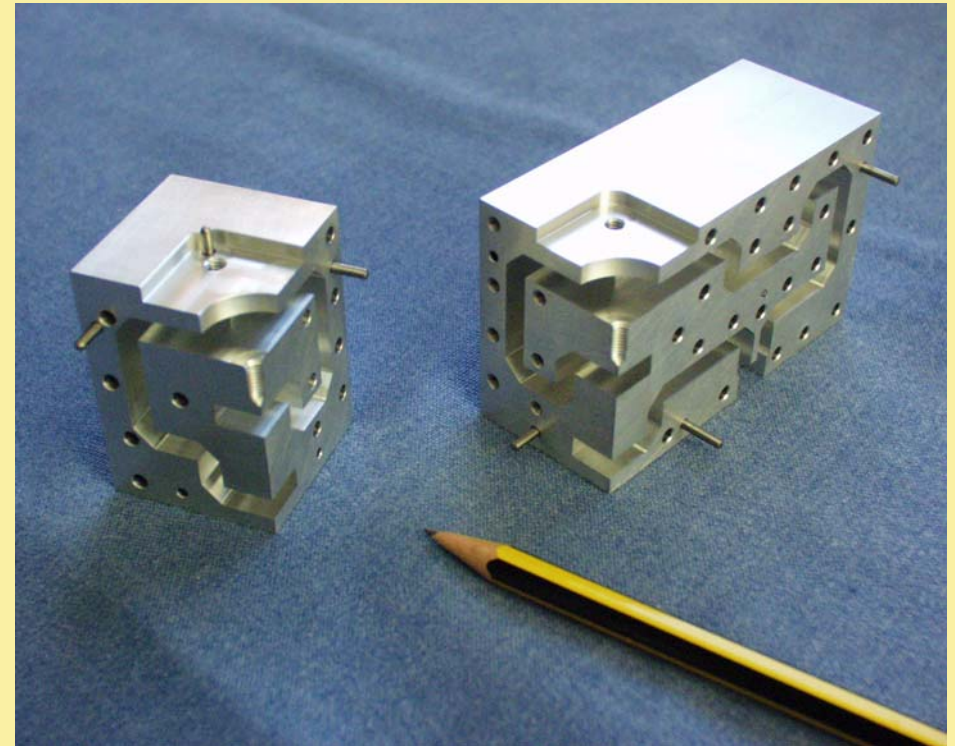
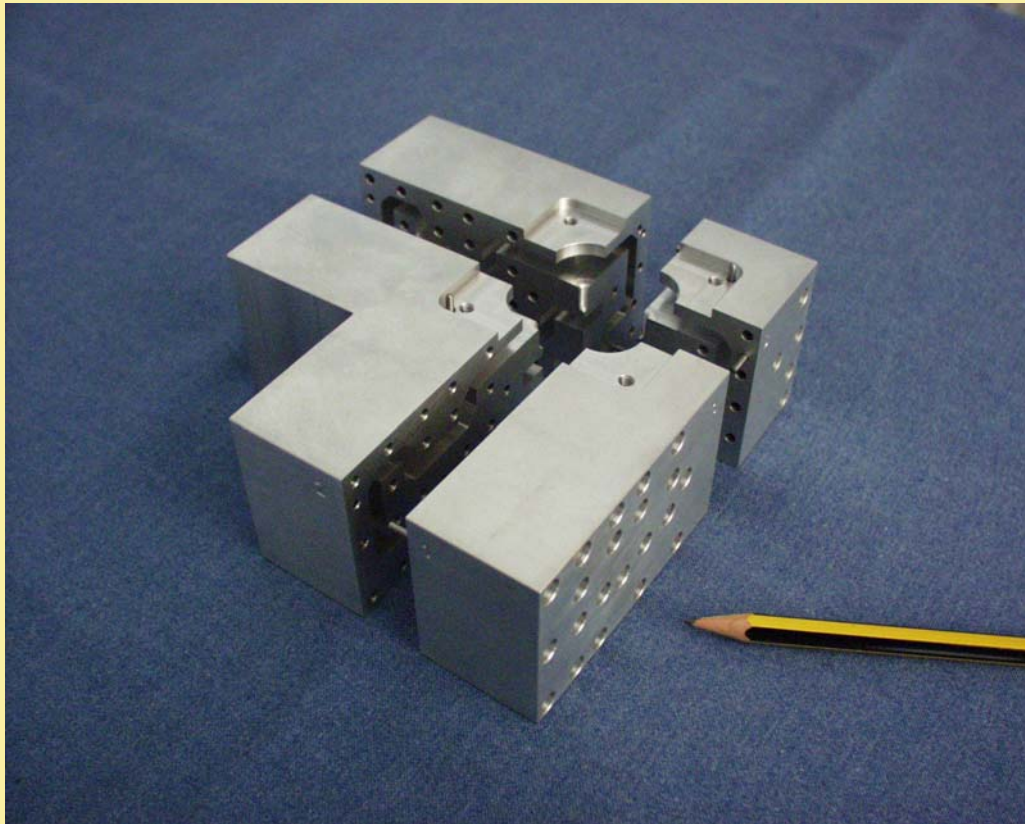
- Separates linear orthogonal polar components
- Sets a limit in the cross-polarization
- Based on turnstile junction
- Phase balanced outputs
- Broadband (> 40 % bandwidth)

Orthomode Transducer (OMT)



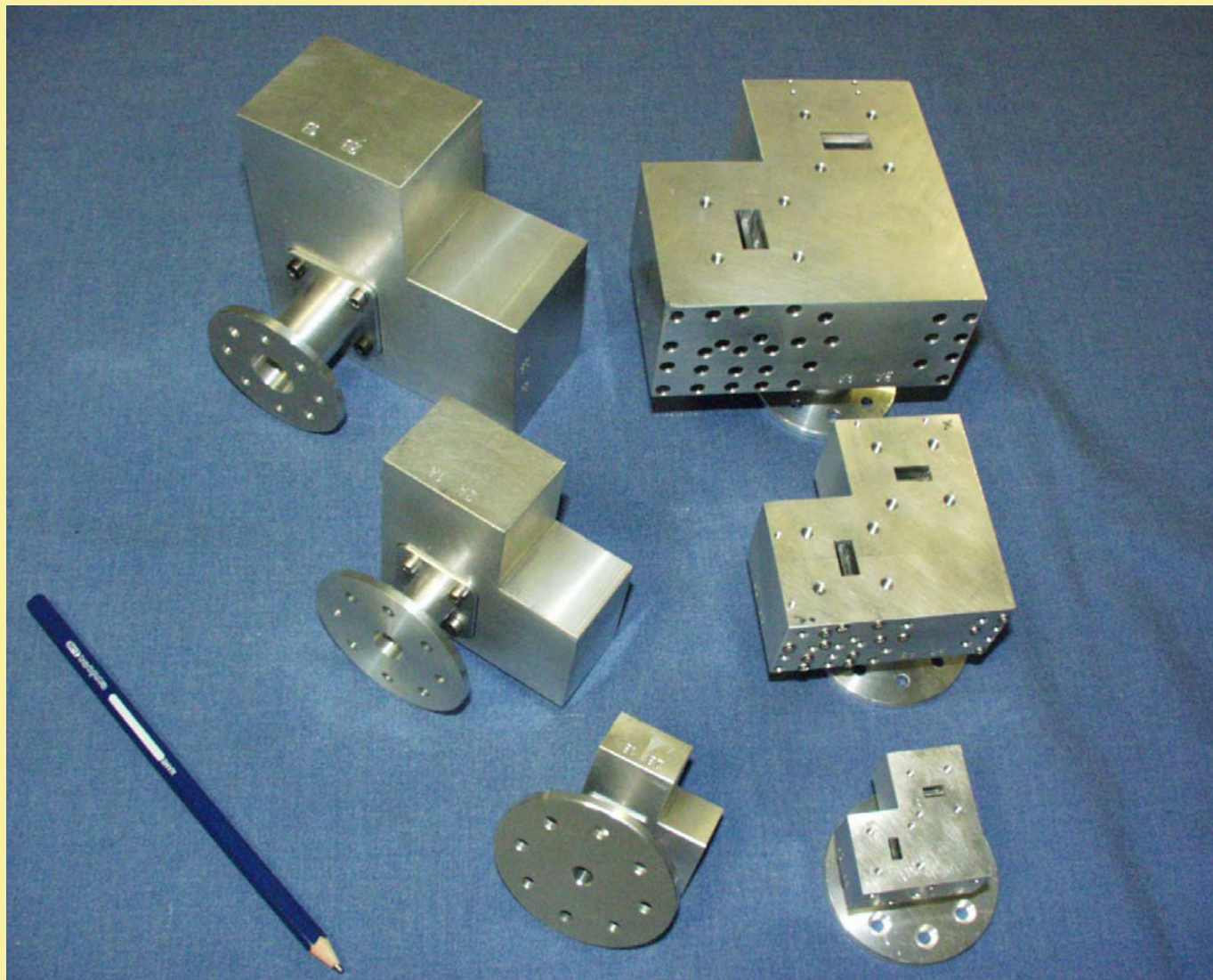
- Reduced height rectangular waveguide
- Optimized E-plane bends
- Scatterer: a critical part of turnstile junction
- Scalable structure (WR75, WR51, WR28)

Orthomode Transducer (OMT)



OMT parts (WR75 version, 10-14 GHz)

Orthomode Transducer (OMT)



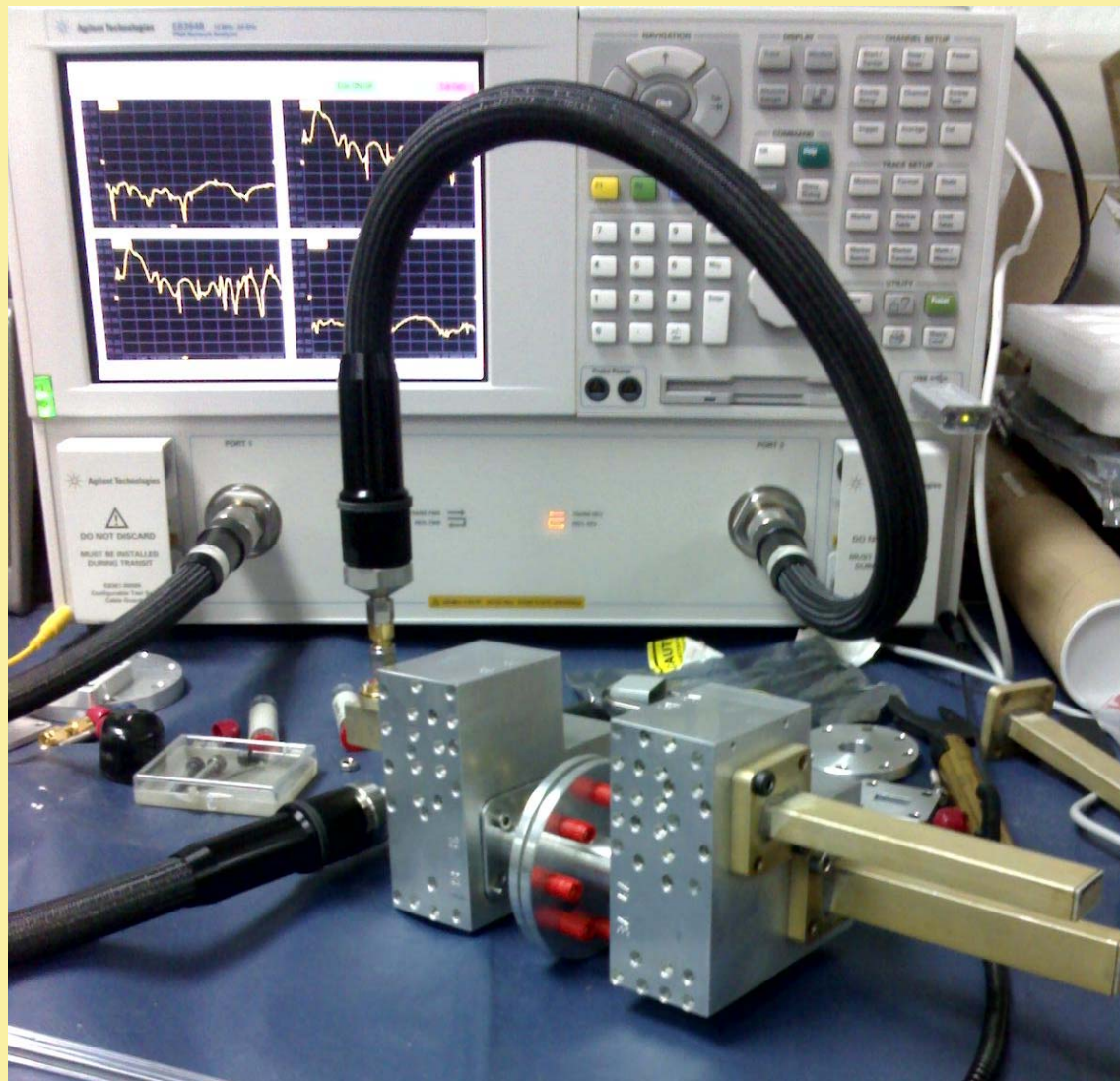
Units:

10-14 GHz

14-22 GHz

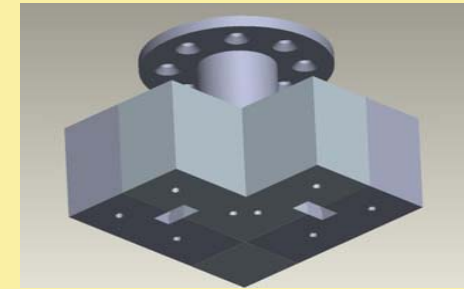
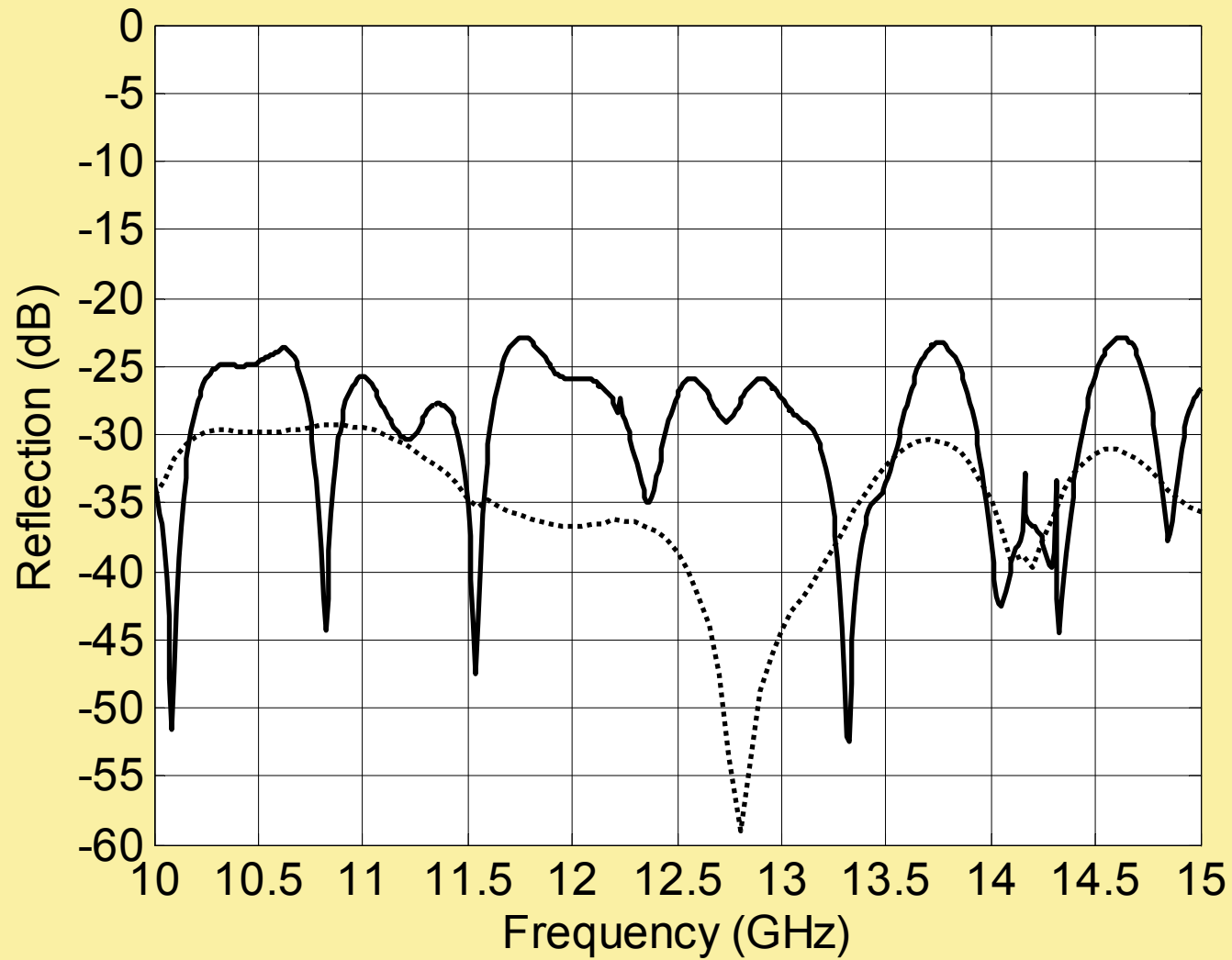
26-36 GHz

Ortho Mode Transducer (OMT) S-parameter tests



- Low return losses
- Low insertion losses
- High isolation
- Excellent phase balance
- Tests at Room Temperature

OMT Return Loss

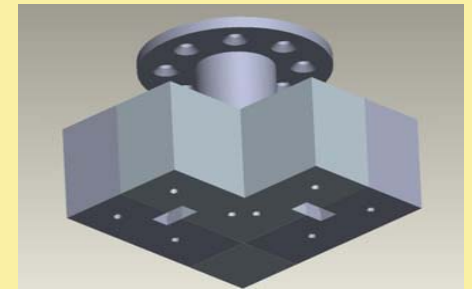
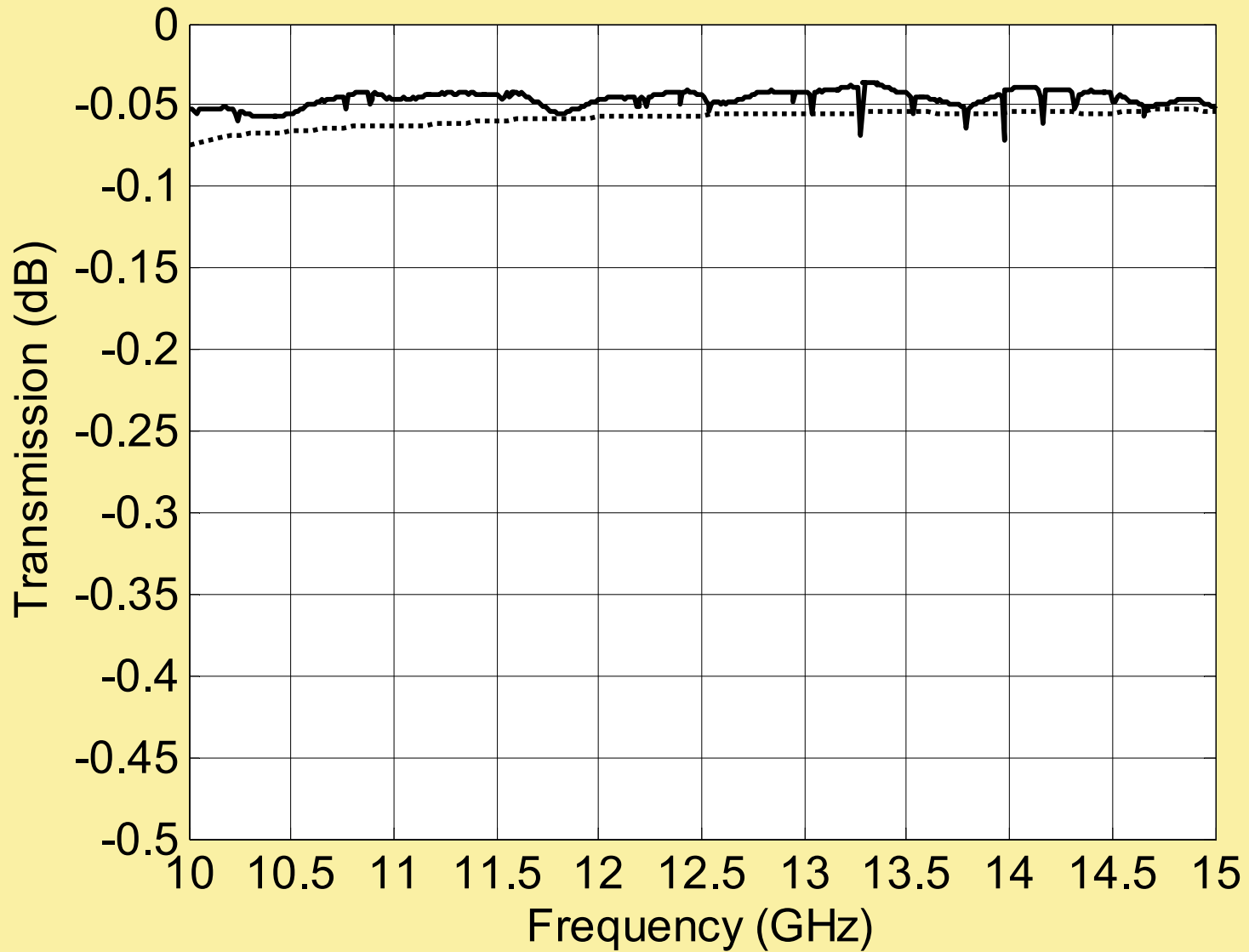


better than 23 dB

solid line: test

dashed line: simulation

OMT Insertion Loss

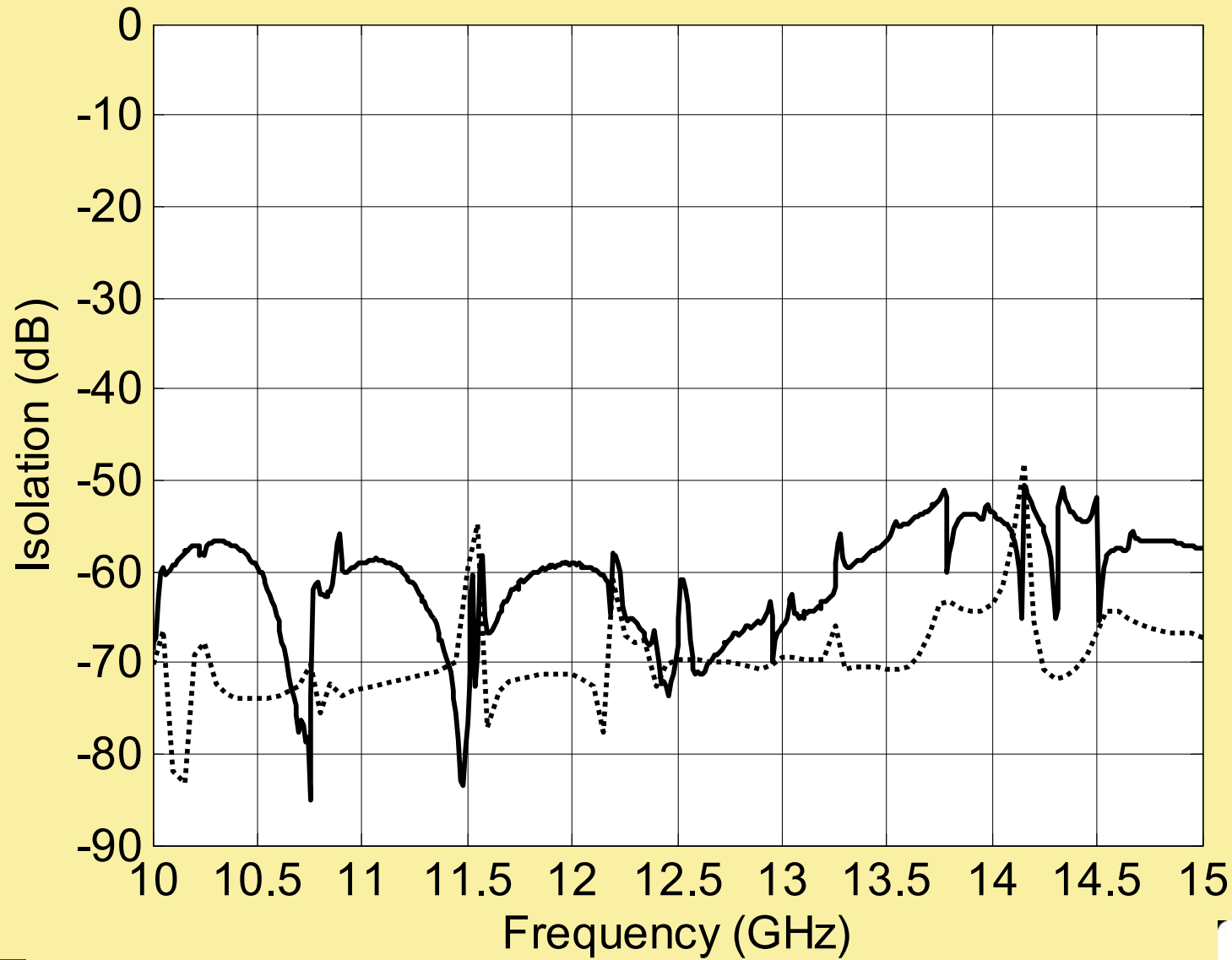
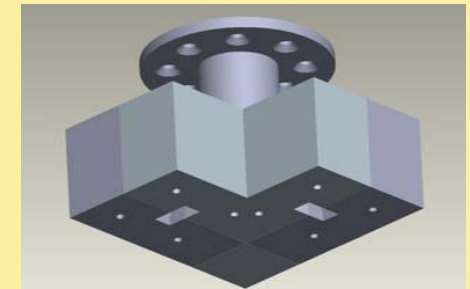


≈ 0.05 dB

solid line: test

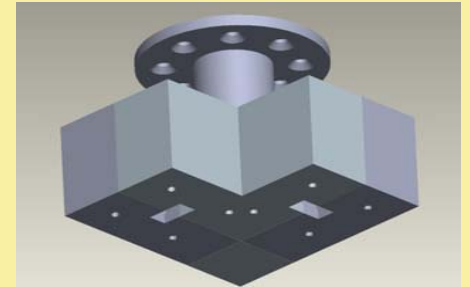
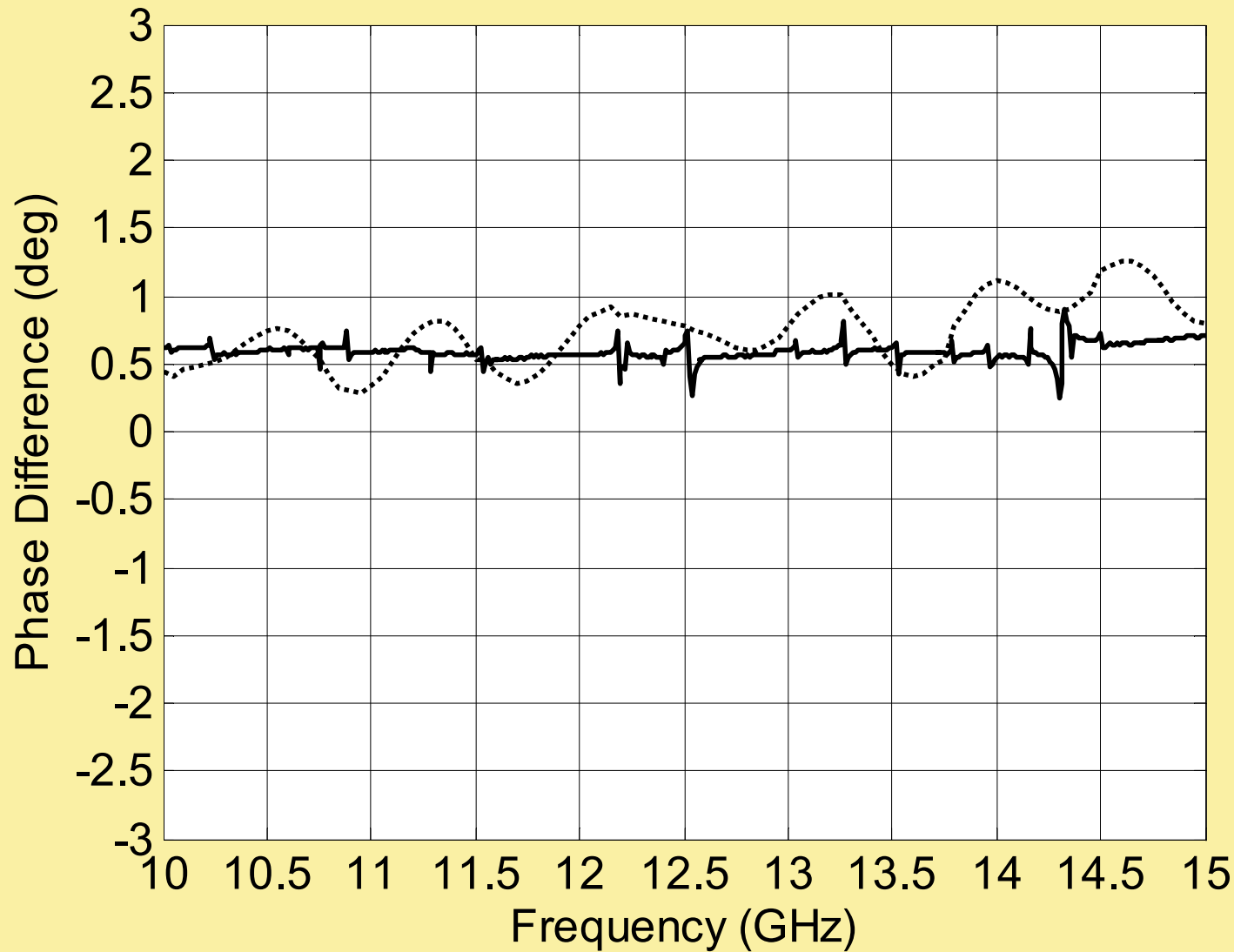
dashed line: simulation

OMT Isolation (between rectangular ports)



better than 50 dB

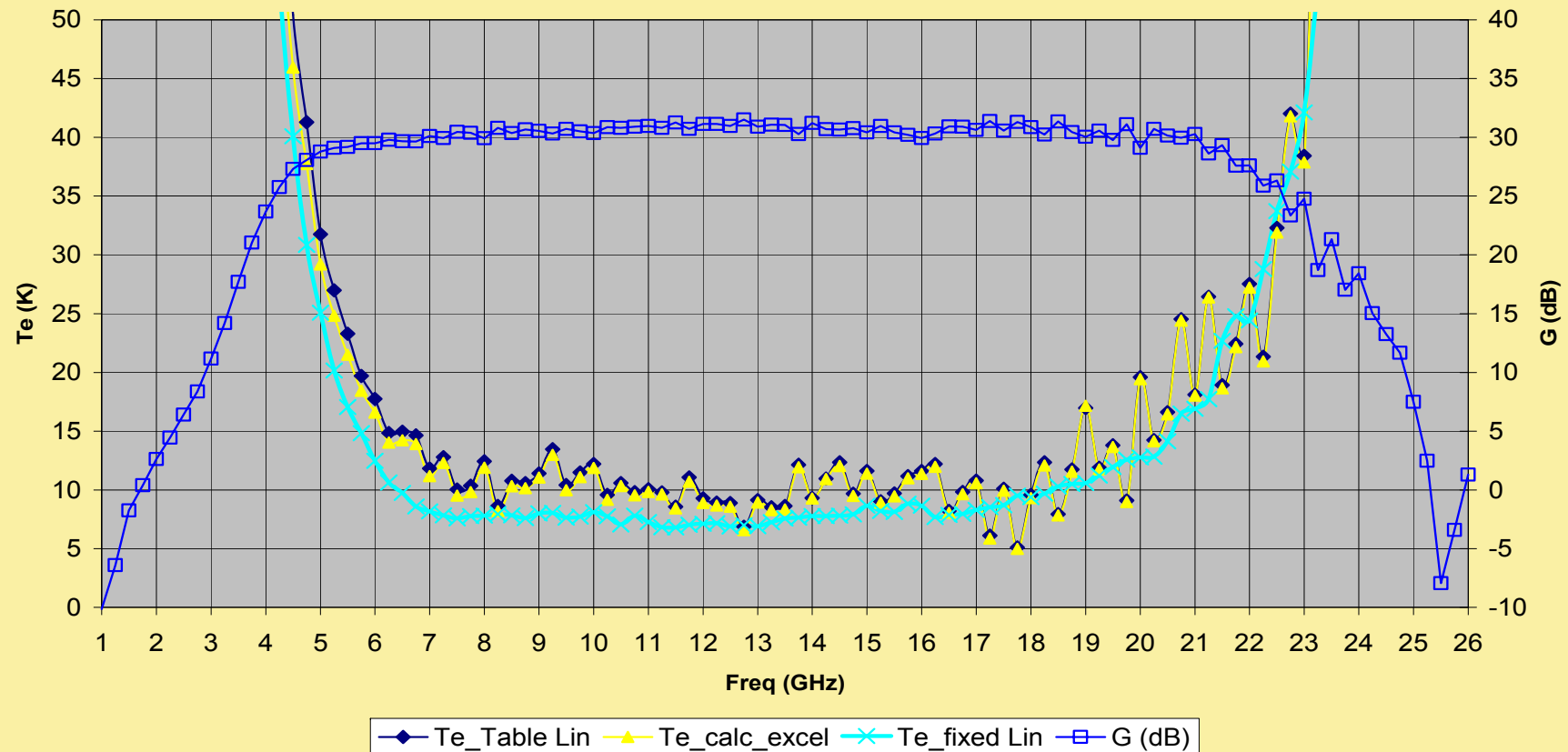
OMT Phase balance (between rectangular outputs)



$\Delta\text{phase} < 0.7^\circ$

Cryo LNA. Gain and Noise Temperature

LNA #40A28 24K 290208
Vg1=1V Vg2=1V Vd=0.65V Id=16mA
NFA 8975A Loss Comp. Table T_input=22.9K



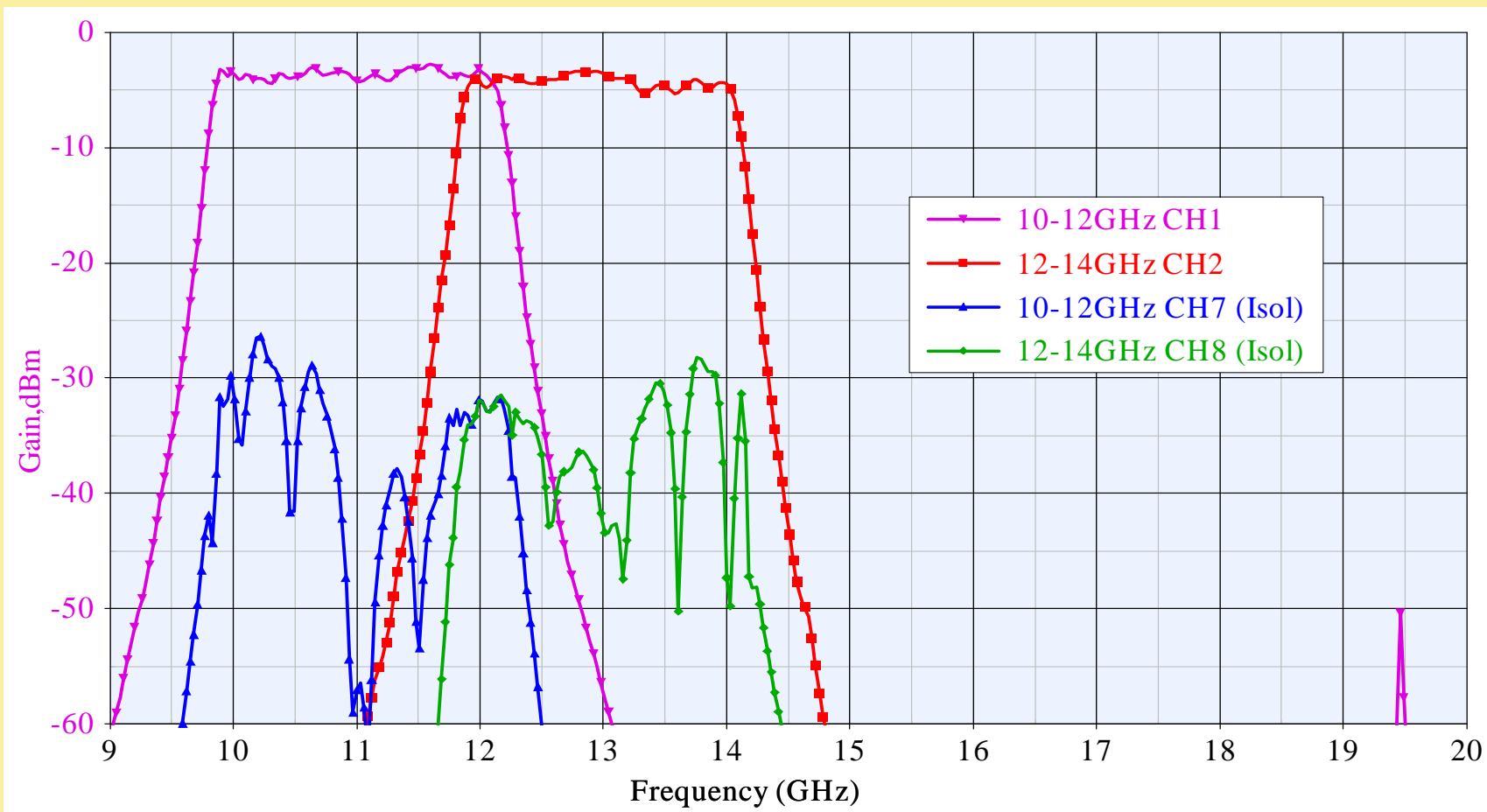
LNA for low frequency channels



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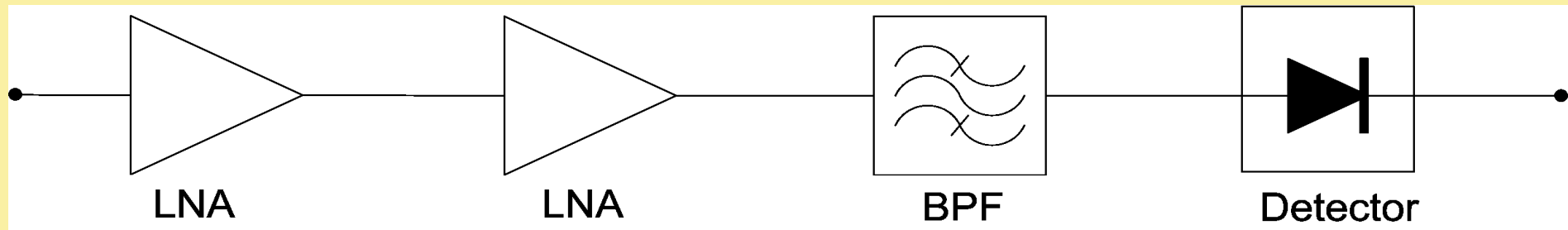


10-14 GHz BEM rack RF Gain (QUIJOTE 1)



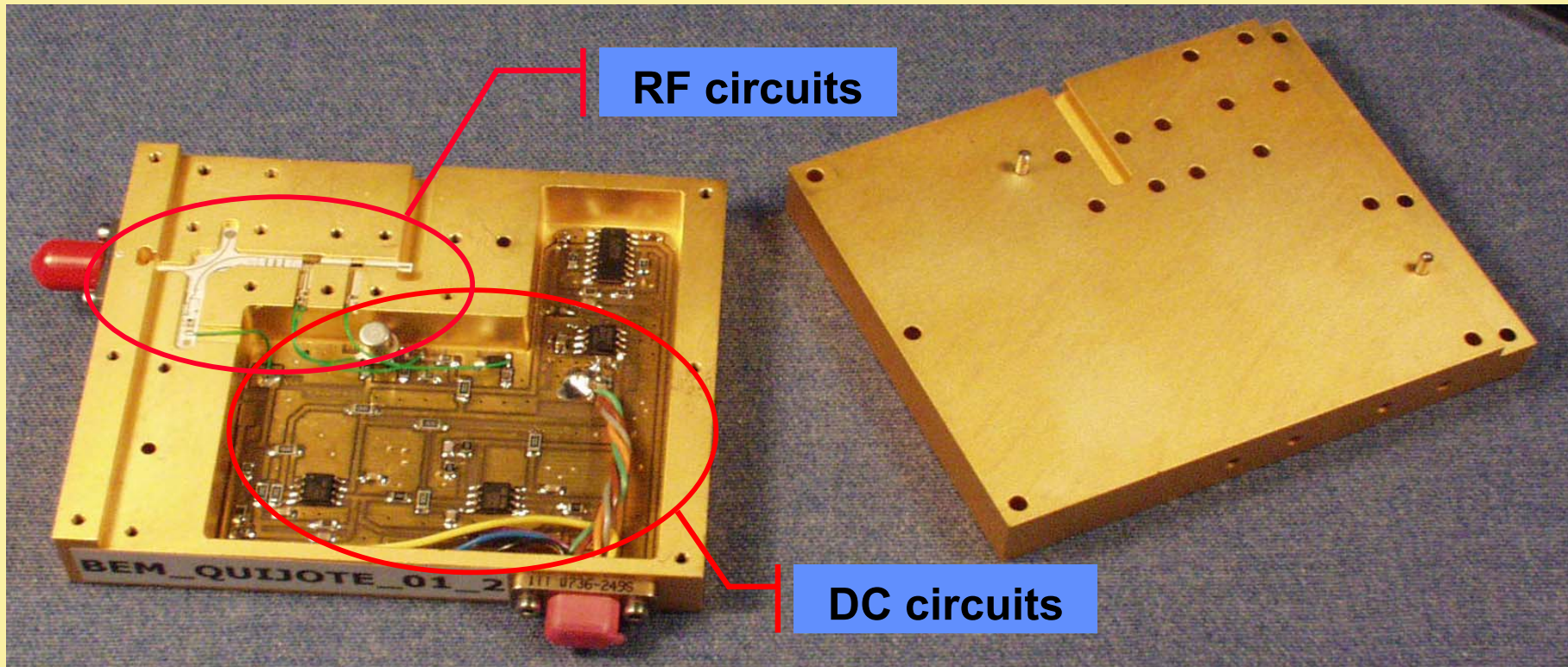
30 GHz BEM (QUIJOTE 1)

30 GHz BEM Block diagram (video amp not included)



Bandwidth: 26 to 36 GHz
Noise temperature < 500 K

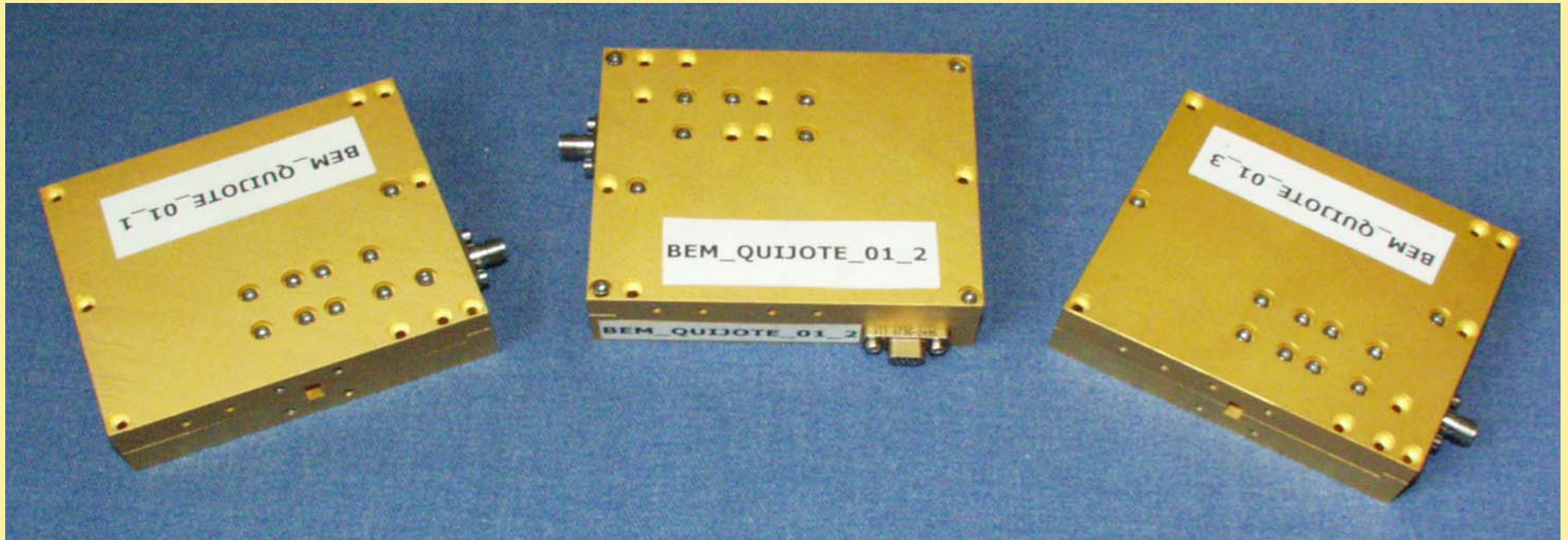
30 GHz BEM (QUIJOTE 1)



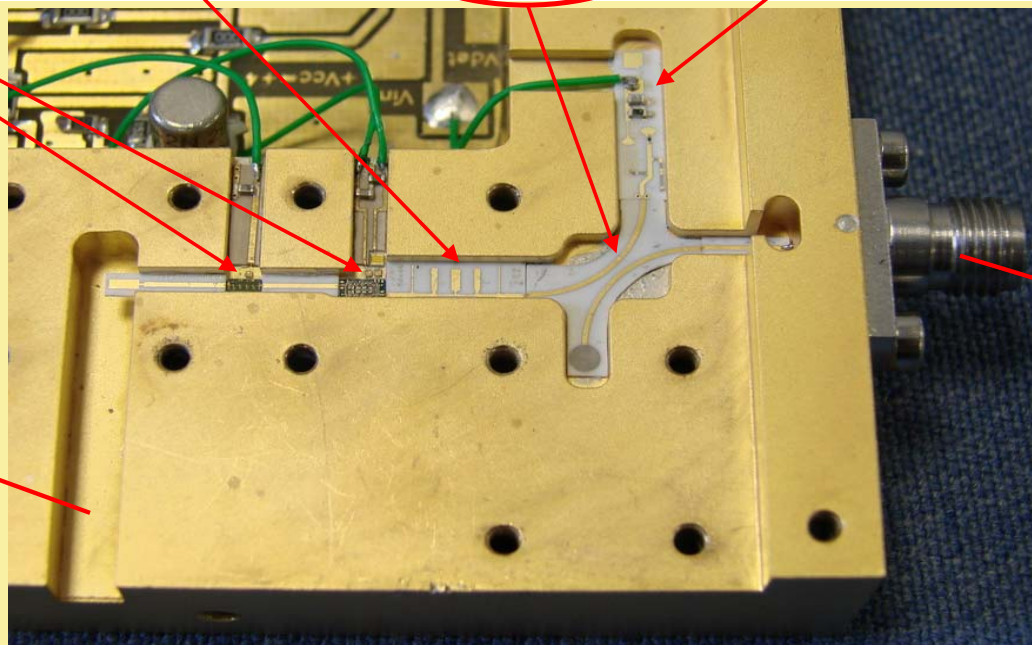
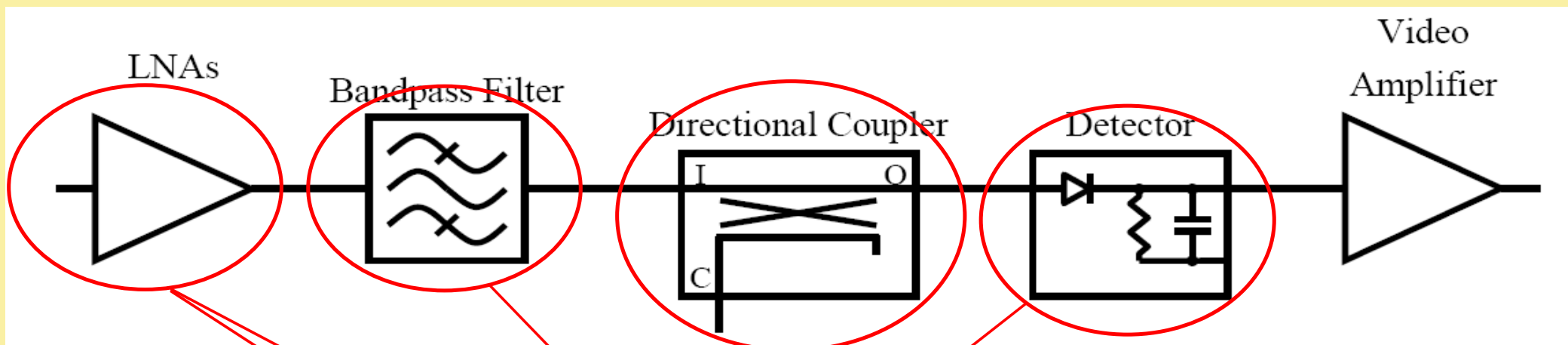
BEM branch (top cover removed)

30 GHz BEM (QUIJOTE 1)

Manufactured BEM units: 2 + 1 (spare)



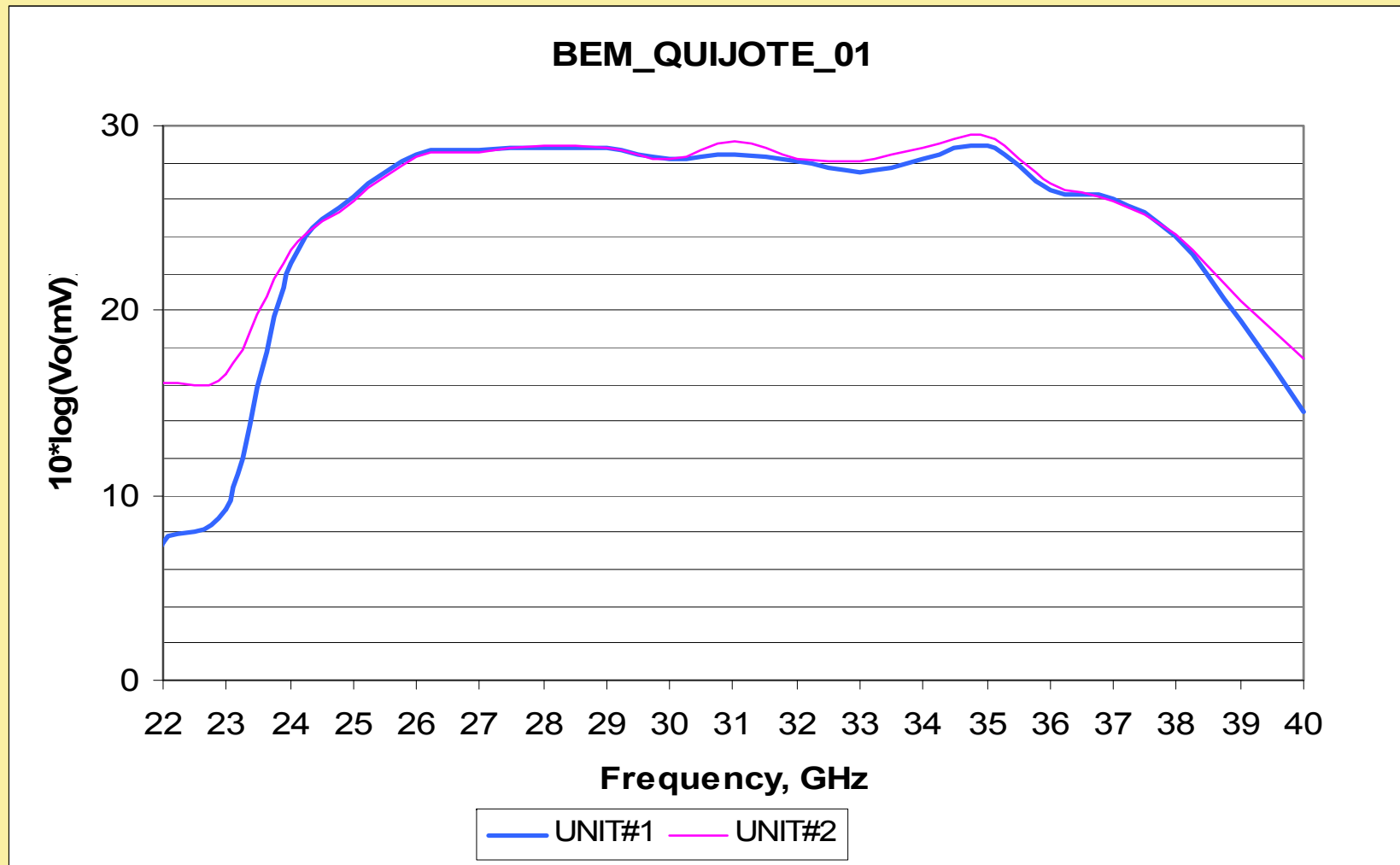
30 GHz BEM (QUIJOTE 1)



Waveguide input

RF sample output (K connector)

Test results (30 GHz BEM)



Gain vs. frequency for two units (detector included)

QUIJOTE 1 - Planning

- Dec. 2009: Cryostat and polar modulators in IAC (Tenerife)
- Jan. 2010: OMT and feed-horns installation
- Mar. 2010: FEM + BEM integration. RF and radiometric testing

Transport telescope to Izaña

- Apr. 2010: Preliminary start-up and commissioning

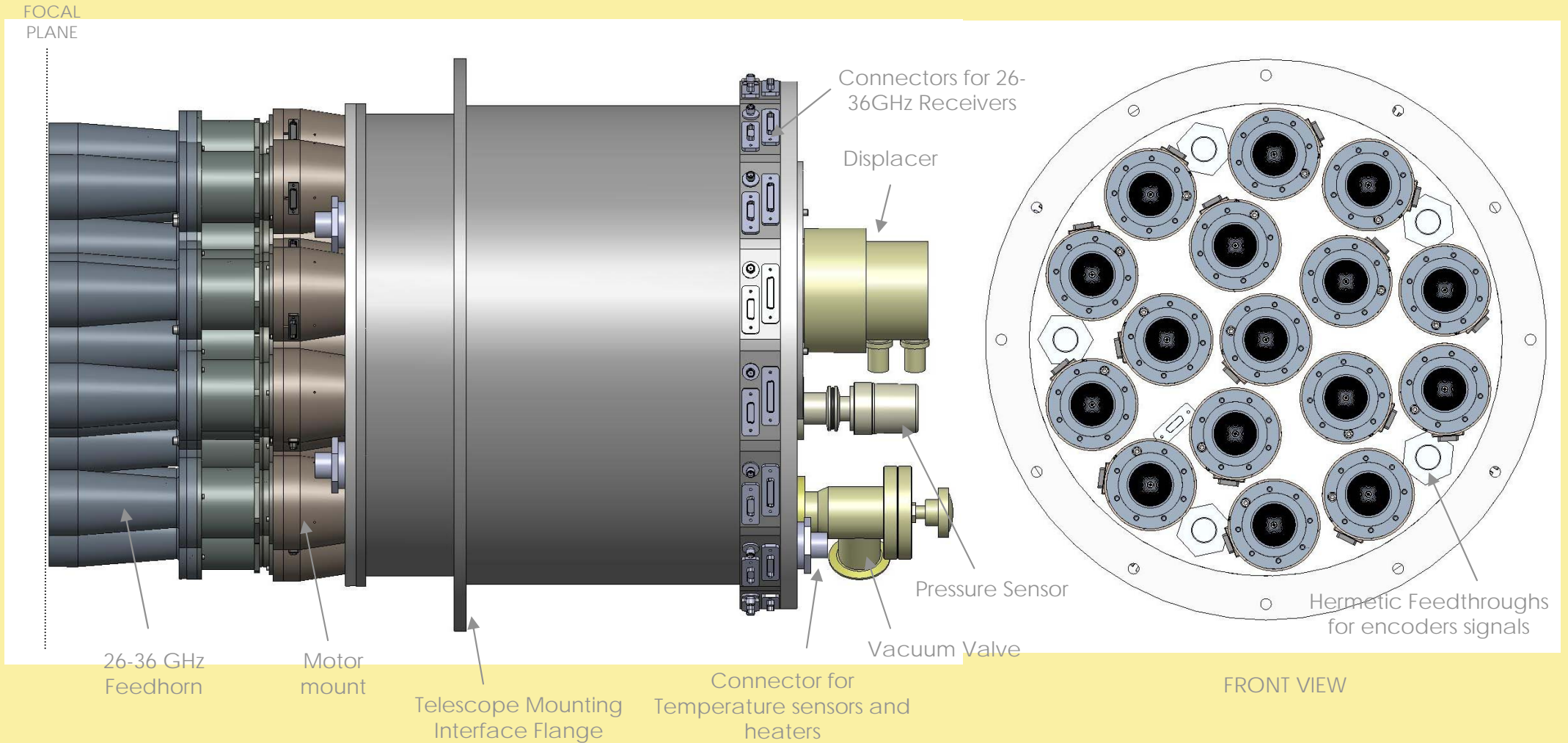


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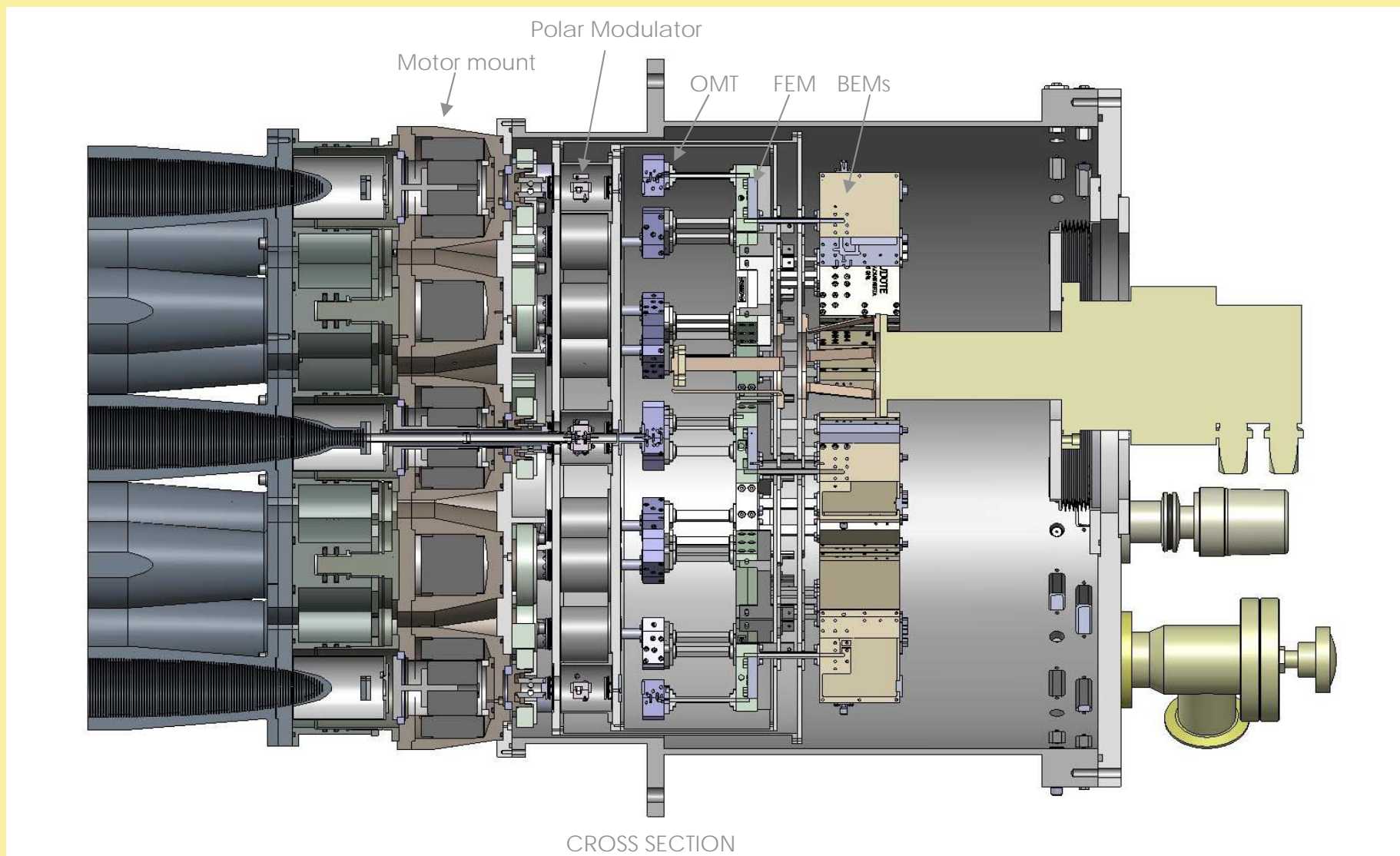


QUIJOTE 2 – 30 GHz Instrument

“If QUIJOTE 1 proves successful”



QUIJOTE 2 – The 30 GHz Instrument



QUIJOTE enclosure building



November 2008: Covering main structure

QUIJOTE enclosure building



May 2009: final quality tests

Telescope manufacturing



**September 2008:
Telescope parts and assembly**

**December 2008:
Telescope shipping**

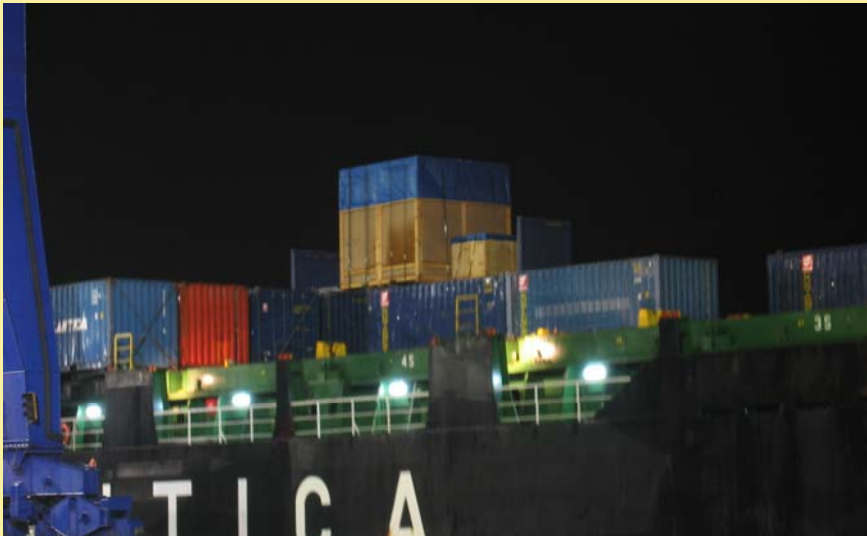


Telescope transport

Dec. 2008 – Jan. 2009: from Bilbao to La Laguna (IAC)



Truck in Bilbao

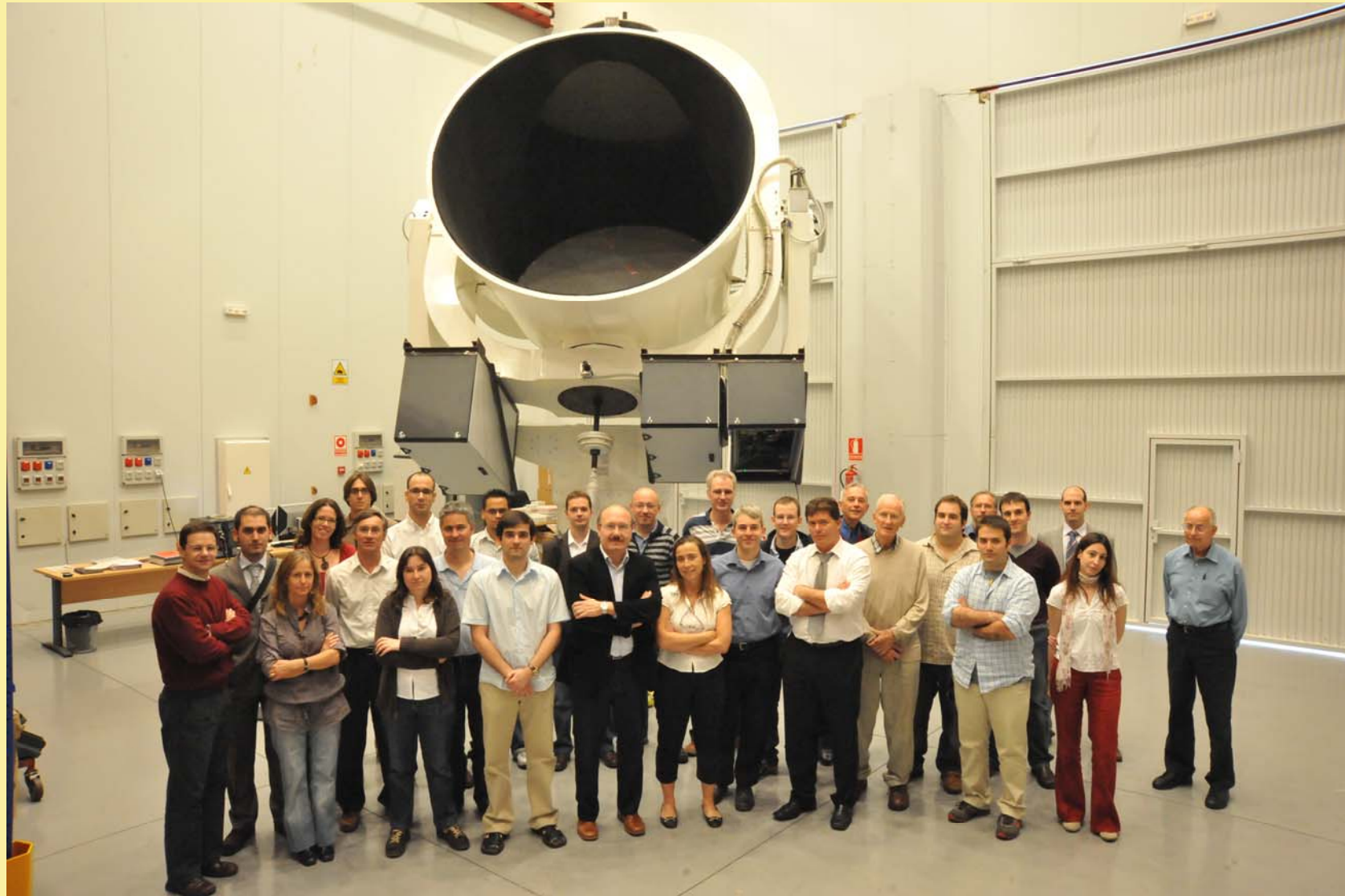


Ship to Tenerife



Telescope QUIJOTE 1 IAC (La Laguna, Tenerife) May 2009

QUIJOTE Second Consortium Meeting (May 2009)



The QUIJOTE CMB Experiment



For further information visit:

<http://www.iac.es/project/cmb/quijote/index.php>