



Attempt at an overview of the CASA "HPC" features mentioning the ALMA pipeline and the ALMA QA2 script generator

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Motivation

CASA Common Astronomy Software Applications

- ALMA datasets are large typical sizes of a single dataset with high spectral res. will be 500 GB (MS format, 0.5 h obs time = 1 SB, 40 antennas)
- CASA (e.g. v3.4, official ALMA version in Cycle 0) is too slow: already processing a ca. 50 GB dataset took ca. 2.5 h on a 3 GHz machine
 => processing one typical SB execution from cycle 1 will take one day at this rate
- Study shows not all processing tasks are I/O limited, especially not when a fast filesystem is available (e.g. SSD or Lustre)
 => can gain from higher computing speed, e.g. by parallelisation
- Up to CASA 3.4, the individual CASA instance runs on a single processor except where OpenMP is used (regridding in imaging, table sorting).
- HPC effort started after CASA 3.4 release. Goal: accelerate CASA by parallelisation



CASA parallelization



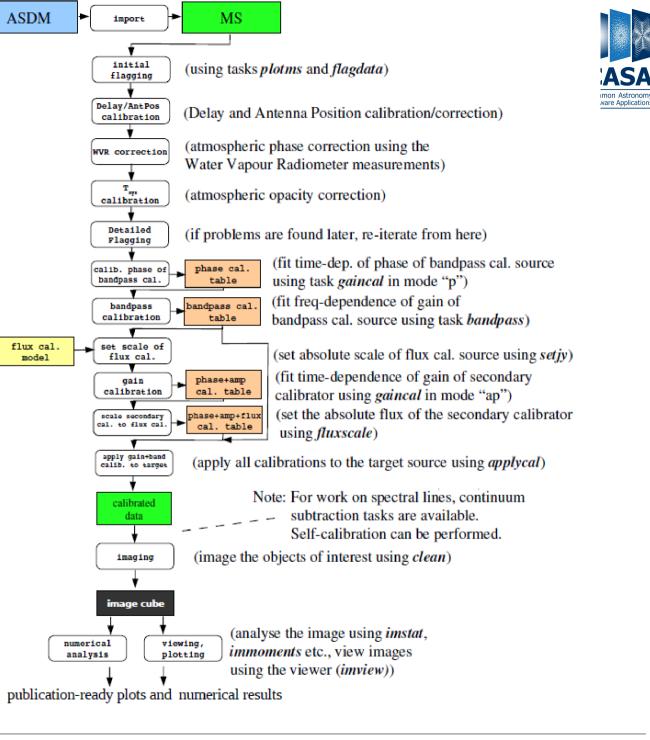
- non-trivial parallelization
 - \rightarrow already implemented in some parts of the imaging code and the table module
 - \rightarrow achieved on the C++ level using OpenMP
 - $\rightarrow\,$ completely transparent to the user
 - $\rightarrow\,$ not applicable to most of the processing tasks
 - \rightarrow time-consuming to develop and debug
- trivial parallelization
 - \rightarrow idea: split the input data into independent chunks, process them in parallel
 - \rightarrow design by J. Kern:
 - task *partition* splits the data along the time and frequency axes into a *Multi-MS (MMS)* - looks like a single MS

but is many sub-MSs underneath

- all subsequent tasks recognize the MMS and *parallelize themselves* to process the sub-MSs accordingly.
- \rightarrow mostly transparent to the user
- \rightarrow applicable to many of the processing tasks, more simple to test and debug
- \rightarrow proven overall speed-up by factors > 2 depending on the file-system used

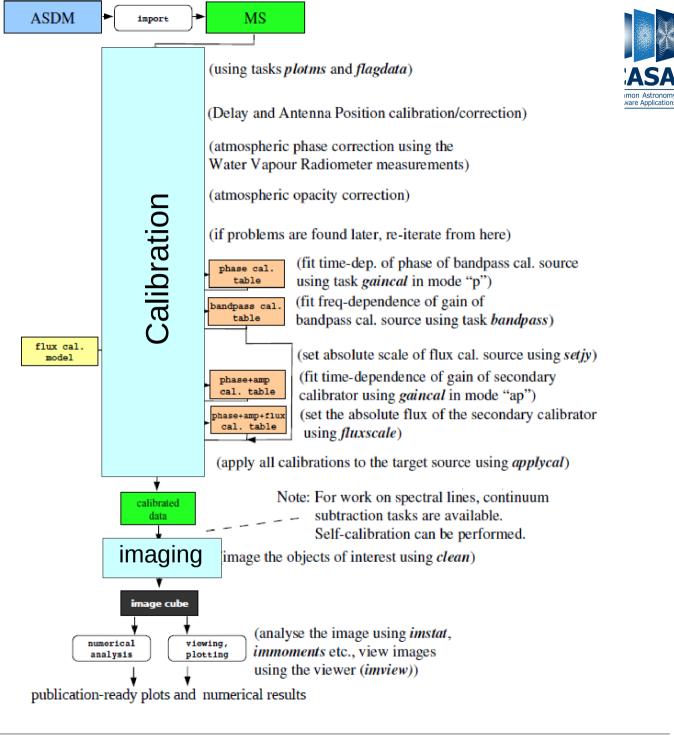


CASA parallelization

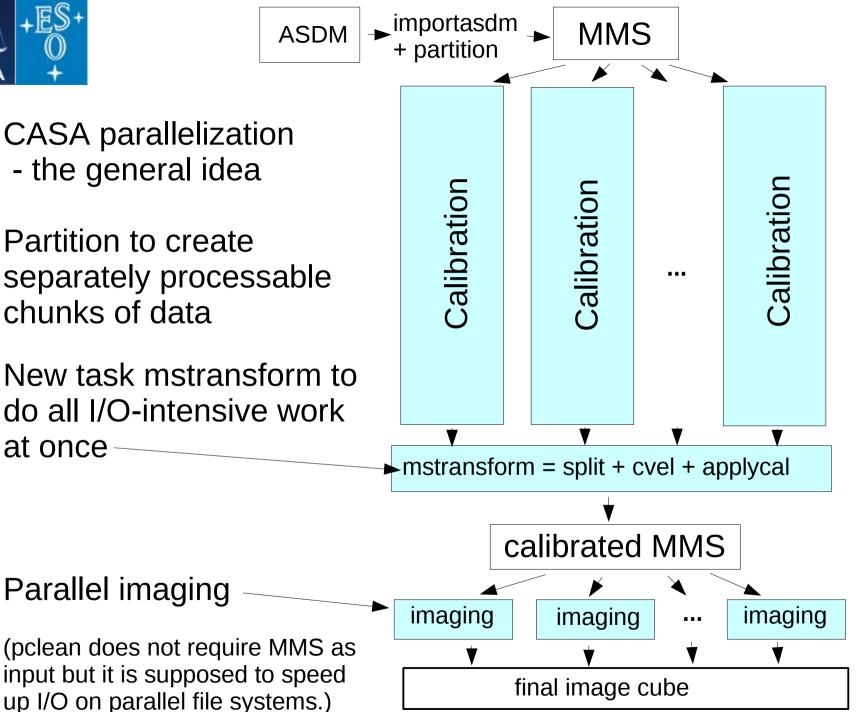




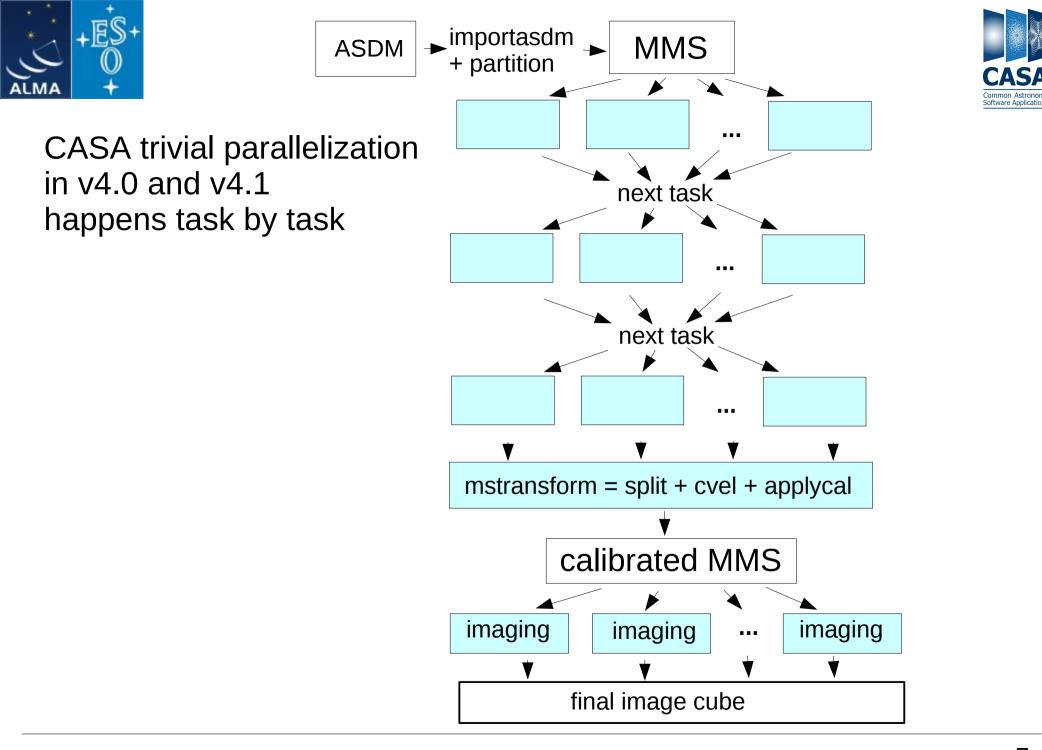
CASA parallelization







D. Petry, Attempt at a summary of CASA "HPC" features, April 2013



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- Work in the CASA 4.0 development cycle (mostly at ESO, May October 2012)
- results:
 - 1: created monitoring and assessment *utilities* for the parallelization framework and the MMSs
 - 2: completed the *partition task* and the *parallelization framework* (created proper unit tests, debugged the existing prototypes, documented)
 - 3: went through nearly all *parallelizable tasks* and made them work in the parallization framework (extended unit tests accordingly, documented), imaging parallelization to be done by K. Golap in Socorro
 - 4: created an *end-to-end regression tests* which uses the new framework and measured speed-up

See new chapter in the CASA 4.0 cookbook.





Results of the 4.0 dev. cycle

- speedup measurements based on ca. 30 GB ALMA datasets were disappointing (only ca. 10-20% speedup overall)
 - presently assume our dataset too small to demonstrate the speedup, but even on a SSD system the speedup is small.
 - Discussion on parallel file systems vs. solid state disks necessary
 - Advantages of a parallel file systems for the individual user running one analysis session still have to be demonstrated.





- Work in the CASA 4.1 development cycle (since end of 2013)
 1) Creation of the "mstransform" task (mostly at ESO),

 a combination of split, cvel, applycal, and partition
 functionality (recurring pattern in analysis scripts)
 to achieve further speed-up by avoiding I/O on intermediate data
 - (at least partially in release 4.1)
 - 2) Ger, Michel, and DP developed "lazy" import in importasdm: new asdmStMan storage manager enables direct access to the ASDM via the DATA column (which is read-only anyway!)
 - => reduces data volume on disk by nearly 50%,
 - speeds up import by ca. 80%,
 - speeds up access (due to compressed storage in ASDM) by 20%,
 - => overall speedup equivalent to elimination of import(!)
 - 3) DP developed "virtualconcat": concatenation of MSs into MMS speedup factor 5 to 10 w.r.t. concat





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The last two points show:

Speed can not only be gained by brute-force parallelisation! Deep understanding of the data structure and clever optimisation of time-consuming processes can give large improvements also without using more than one processor!



The ALMA Pipeline

CASA Common Astronomy Software Applications

- No time to go into the decade-long history of the ALMA pipeline
- Status now:
 - calibration section nearing completion for cycle 0 features
 - imaging section still rudimentary
- Ideas:
 - Separated Heuristics and Context routines
 - Analysis in "stages"
 - User submits "Pipeline Processing Request" (PPR), pipeline does the rest and takes all decisions automatically including reiterations etc.
 - To be embedded in CASA and distributed with it, anybody can rerun it
- Expect to switch ALMA QA2 calibration to pipeline late this year





- For commissioning and quality assurance, JAO and the ARCs developed the *QA2 Script Generator* (lead: E. Villard)
- general idea: given an raw ALMA MS, the generator creates analysis scripts which do the entire data analysis
- separate script for calibration and imaging
- scripts meant to be run by the QA2 expert analysts AND the science end user
- initially the expert analysis runs the scripts step by step (special stepping mechanism implemented) and makes corrections if necessary (e.g. additional flagging)
- when the script is OK, it can be run in one go or in groups of steps
- final scripts are part of the delivery to the science end user